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36 **1. Introduction**

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38 The mineral strashimirite $Cu_8(AsO_4)_4(OH)_4$:5H₂O is a hydrated hydroxyl divalent 39 copper arsenate[1, 2]. The mineral is of monoclinic symmetry with space group 40 *P*21/*m* and may be compared with other hydroxyl copper arsenate minerals, including 41 euchroite Cu₂(AsO₄)(OH)^{3H₂O, olivenite Cu₂(AsO₄)(OH), cornwallite} $\text{Cu}_5(\text{AsO}_4)_{2}(\text{OH})_4$. Olivenite is monoclinic, space group $P2_1/n$ and is the most 43 common secondary mineral of the oxidized zone of hydrothermal deposits. Other 44 related copper and arsenate minerals are the minerals cornwallite $\text{[Cu}_5(\text{AsO}_4)_2(\text{OH})_4$ 45 and clinoclase $[Cu_3(AsO_4)(OH)_3]$. Each of these minerals occurs in the oxidized 46 zones of copper deposits and olivenite is by far the commonest [3, 4]. Cornwallite and 47 clinoclase are rare secondary mineral that crystallise monoclinic, space group *P*21/*a* 48 [3, 5, 6].

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50 The relative stabilities of the basic copper arsenates have been determined 51 using estimated chemical parameters and experimentally determined solubility 52 products [7, 8]. Magalhaes *et al.* have reported the relative stability of copper 53 arsenate minerals [9]. Normal anhydrous copper(II) arsenate is known as the 54 naturally occurring species lammerite $[C_{u3}(AsO₄)₂]$, but it is very rare [10]. The more 55 basic stoichiometries occupy fields at higher pH as expected. Since olivenite is the 56 stable phase under chemical conditions intermediate to those that serve to stabilize 57 cornwallite and clinoclase, paragenetic relationships have been explored. [10] 58 Olivenite is often found with either cornwallite or clinoclase, but not together. The 59 stability of the basic copper arsenate minerals is related to their redox potential and 60 phase fields exist for the related minerals olivenite, cornubite $\text{[Cu}_5(\text{AsO}_4)_2(\text{OH})_4$], 61 clinoclase and cornwallite. Thus, the structure of these phase related minerals are 62 related and should provide similar spectra, which should differ in terms of the 63 intensity and position of the bands according to the relative mole ratios of Cu/As/OH 64 and the number and site symmetry of the formula units in the unit cell of the minerals. 65

- 96 The minerals were analysed by electron microprobe (Cameca SX100, WD 97 mode) for chemical composition. The results (CR - mean of 6 point analysis) are CaO 98 0.51, FeO 0.06, CuO 50.31, CoO 0.30, NiO 0.54, ZnO 0.38, Al₂O₃ 0.06, Sb₂O₃ 0.08, 99 SiO₂ 0.09, As₂O₅ 37.20, P₂O₅ 0.04, SO₃ 0.55, H₂O_{calc.} 10.44, sum 100.57 wt. % 100 resulting to empirical formula $\left(Cu_{7.59}Ca_{0.11}Ni_{0.09}Zn_{0.05}Co_{0.04}Al_{0.02}Fe_{0.01}\right)_{\Sigma7.91}$ 101 $[(AsO₄)_{3.89}(SO₄)_{0.08}(SiO₄)_{0.02}(PO₄)_{0.01}]_{54.00}(OH)_{3.92}$. 5H₂O. For SR sample (mean of 2 102 point analysis) are CaO 0.06, CuO 50.05, CoO 0.02, NiO 1.72, ZnO 0.22, Al₂O₃ 0.17, 103 Sb2O3 0.02, SiO2 0.09, As2O5 35.53, P2O5 0.88, SO3 0.17, H2Ocalc. 10.46, sum 99.39 104 wt. % and empirical formula $(Cu_{7.74}Ni_{0.28}Al_{0.04}Zn_{0.03}Ca_{0.01})_{\Sigma8.10}$ 105 $[(AsO₄)_{3.80}(PO₄)_{0.15}(SO₄)_{0.02}(SiO₄)_{0.02}]_{23.99}(OH)_{4.27} . 5H₂O. The basis of recalculating$ 106 is $(As+Si+P+Sb) = 4$ apfu; water contents were calculated on the basis of charge 107 balance and theoretical content of $H_2O = 5$ pfu. The chemical analysis of the two 108 strashimirite minerals are given in Table 2. The composition of this mineral has been 109 reported by Anthony *et al.* [29] (page 565). 110 111 The sample of cornubite from the Daly mine, Flinders Ranges, South Australia 112 was obtained from the collections of the Mineralogy Section, Museum Victoria, 113 Melbourne, Victoria, Australia (specimen M37793). Cornwallite and olivenite from 114 the Penberthy Croft mine, St Hilary, Cornwall, UK, were supplied by Mr John 115 Betterton. Samples of olivenite and clinoclase from the Tin Stope, Majuba Hill mine, 116 Utah, USA, were purchased from the Mineralogical Research Company. All were 117 checked for purity by powder X-ray diffraction and by SEM and microprobe methods. 118 Negligible amounts of phosphorus or transition metals other than copper were found 119 in the samples used for this spectroscopic study. 120 121 *2.2. Raman spectroscopy* 122 123 The crystals of strashimirite were placed on the stage of an Olympus BHSM 124 microscope, equipped with 10x and 50x objectives and part of a Renishaw 1000 125 Raman microscope system, which also includes a monochromator, a filter system and 126 a Charge Coupled Device (CCD). Raman spectra were excited by a HeNe laser (633 127 mm) at a resolution of 2 cm⁻¹ in the range between 100 and 4000
	- 4

128 cm⁻¹. Repeated acquisition using the highest magnification was accumulated for 64 129 scans at 20 secs. to improve the signal to noise ratio. The instrument was calibrated 130 using the 520.7 cm^{-1} line of a silicon wafer.

131

132 *2.3 Infrared spectroscopy*

133 The FTIR spectrum of strashimirite was obtained with the FTIR Nicolet 740 134 spectrometer using the conventional KBr-disk technique. Infrared spectrum in the 135 range 4000-400 cm⁻¹ was obtained by the co-addition of 32 scans with a resolution of 2 cm^{-1} and a mirror velocity of 0.1496 cm/s. Spectral manipulation such as baseline 137 adjustment, smoothing and normalization were performed using the OMNIC software 138 package (Thermo Electron Corporation). Band component analysis was undertaken 139 using the same software package which enabled the type of fitting function to be 140 selected and allows specific parameters to be fixed of varied accordingly. Band fitting 141 was done using a Lorentz-Gauss cross-product function with the minimum number of 142 component bands used for the fitting process.

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145 Spectral manipulation such as baseline adjustment, smoothing and 146 normalisation were performed using the Spectracalc software package GRAMS 147 (Galactic Industries Corporation, NH, USA). Band component analysis was 148 undertaken using the Jandel 'Peakfit' software package which enabled the type of 149 fitting function to be selected and allows specific parameters to be fixed or varied 150 accordingly. Band fitting was done using a Lorentz-Gauss cross-product function with 151 the minimum number of component bands used for the fitting process. The Gauss-152 Lorentz ratio was maintained at values greater than 0.7 and fitting was undertaken 153 until reproducible results were obtained $(r^2 > 0.995)$. 154 155 156 157 **3. Results and discussion** 158

159 *3.1. Raman spectroscopy*

161 The Raman spectra of the tetrahedral anions in aqueous systems are well 162 known [30]. The Raman active symmetric stretching v_1 vibration of the arsenate 163 anion is observed at 810 cm^{-1} and coincides with the position of the triply degenerate 164 Raman and infrared active antisymmetric stretching mode (v_3) . The doubly 165 degenerate Raman active symmetric bending mode (v_2) is observed at 342 cm⁻¹ and 166 the triply degenerate Raman and infrared active out-of-plane bending modes (v_4) is 167 observed at 398 cm⁻¹. Of all the tetrahedral oxyanions spectra, the positions of the 168 arsenate vibrations occur at lower wavenumbers for naturally occurring mineral 169 oxyanions spectra [30]. Farmer lists a number of infrared spectra of arsenates 170 including the basic copper arsenates olivenite and euchroite[31]. Depending on the 171 symmetry of the anion, the bands may be infrared active but not Raman active, but 172 may be both. The effect of the arsenate ion in a crystal will be to remove the 173 degeneracy and allow splitting of the bands according to factor group analysis. The v_1 174 and v_3 bands of olivenite and euchroite were observed at 860, 828, 790 cm⁻¹ and at 175 830 and 770 cm⁻¹ respectively. The $v_4 (AsO₄³)$ bending modes were found at 493 176 and 452 cm⁻¹ for olivenite and at 475 and 410 cm⁻¹ for euchroite. No v_2 bands were 177 registered. This is no doubt related to the fact the bands are found below 400 cm⁻¹. 178 Two OH stretching vibrations were observed in the infrared spectrum at 3580 and 179 3440 cm⁻¹ for olivenite. The Raman and infrared spectra of strashimirite minerals A 180 and B are shown in Figures 1 and 2 respectively. The results of the Raman and 181 infrared spectra of strashimirite minerals together with the results of the Raman 182 spectra of olivenite, cornwallite, clinoclase and cornubite are reported in Table 3. 183

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185 The Raman spectra of strashimirite from the two different deposits in the 186 region 700 to 1000 cm⁻¹ are shown in Fig. 3. In the Raman spectrum of strashimirite 187 A, a broad intense band centred at 850 cm^{-1} is observed and may be band component 188 analysed into components at 831 and 852 cm⁻¹ which are assigned to the $v_1 (AsO₄³)$ 189 symmetric and v_3 (AsO₄³) antisymmetric stretching modes. Low intensity bands are 190 also observed at 893 and 982 cm^{-1} . The latter is attributed to some phosphate 191 substitution, as is confirmed by X-ray diffraction and infrared spectroscopy. A 192 comparison may be therefore made with the infrared data of this mineral. In the

- 193 infrared spectrum of A (Fig. 4) intense infrared bands are observed at 760 and 833
- 194 cm⁻¹ with two shoulder bands at 804 and 909 cm⁻¹. The latter band at 909 cm⁻¹
- 195 corresponds reasonably well with the Raman band at 893 cm⁻¹ and the 833 cm⁻¹
- 196 infrared band corresponds to the Raman band at 839 cm^{-1} . Low intensity infrared
- 197 bands are observed at 1002, 1085 and 1123 cm⁻¹ which are assigned to $(PO₄³)$ for
- 198 (AsO₄³) substitution or to (SO_4^2) substitution.
- 199

200 The Raman spectrum of strashimirite B in the 700 to 1000 cm⁻¹ region shows 201 greater complexity. Bands are observed at 826, 843, 891 and 988 cm^{-1} . The band at 202 988 cm⁻¹ may be probably attributed to the $v_1 (PO_4^3)$ or $v_1 (SO_4^2)$. A band is also 203 observed at 771 cm⁻¹. The attribution of which is unknown. One possibility is that 204 this is a water librational mode. Such a band should show greater intensity in the 205 infrared spectrum but this is not the case. The two bands at 826 and 843 cm^{-1} are 206 assigned to the $v_1 (AsO_4^3)$ symmetric and $v_3 (AsO_4^3)$ antisymmetric stretching 207 modes. The infrared spectrum of strashimirite B is reported in Fig. 1b (supplementary 208 information). An intense infrared band is found at 843 cm^{-1} with two shoulder bands 209 at 826 and 891 cm⁻¹. The correspondence between the infrared data and Raman data 210 is good for strashimirite B. Because of the diagenetic relationship between the basic 211 copper arsenate minerals, a comparison of the Raman spectra with the other minerals 212 can be made (see Table 3). The Raman spectrum of olivenite shows two $(AsO₄³)$ 213 bands at 853 and 820 cm⁻¹; cornwallite at 859 and 806 cm⁻¹, cornubite at 815 and 780 214 cm⁻¹ and clinoclase at 823 and 771 cm⁻¹[30, 32]. Additional very weak Raman bands 215 for olivenite are observed at 880 and 790 cm^{-1} . The most intense band is assigned to 216 the v_1 symmetric stretching vibration. This assignment differs from that described by 217 Sumin de Portilla [25]. In his work, the v_3 mode was described as splitting into four 218 components at 870, 830, 800 and 750 cm⁻¹. Farmer suggested that the v_1 and v_3 219 modes overlapped and were to be found at the same frequency [31]. Whilst this is 220 highly unusual, it is suggested that the two vibrations at 853 and 820 cm⁻¹ are the $(AsO₄³)$ symmetric and antisymmetric stretching vibrations, respectively. Griffith 222 reported the Raman spectrum of olivenite^[33] at 880 $(A¹)$, 856 (B_{2u}) , 810 $(A¹)$ and 790 223 (B_{2u}) cm⁻¹. The observation of the bands 853 and 820 cm⁻¹ is in good agreement with 224 the data published by Griffith [33]. The most intense bands in the Raman spectra are 225 the bands at 853 and 810 cm⁻¹. Factor group analysis suggests that there should be

- 226 one active Raman band and one active infrared band in the $(AsO₄³)$ stretching region. 227 Additional bands result from a loss of site symmetry. The difference in intensity is 228 related to the number of $(AsO₄³)$ units involved in this site symmetry reduction. Thus 229 the other two infrared bands observed at 800 and 750 cm⁻¹ for olivenite are the 230 corresponding v_3 vibrations. The second band observed at 806 and 771 cm⁻¹ is more 231 intense for cornwallite and clinoclase. The $(AsO₄³)$ stretching vibration for olivenite 232 and cornwallite are in similar band positions, suggesting a similar molecular structure. 233 For cornubite two bands are observed at 815 and 780
- 234 cm^{-1} .
- 235

236 Stoilova and Minceva-Stefanova reported the infrared spectrum of a type 237 strashimirite mineral from Bulgaria[1]. These researchers identified infrared bands at 238 766 and 863 cm⁻¹ and assigned these bands to the v_1 (AsO₄³) symmetric and v_3 239 antisymmetric stretching modes. The intense Raman bands at 845 cm^{-1} for 240 strashimirite A and 843 cm⁻¹ for strashimirite B tend to suggest that the assignation of 241 Stoilova and Minceva-Stefanova may compliment the Raman and infrared results 242 reported in this work. . Further the bands are highly polarised, confirming the 243 assignment of the Raman band. Magalhaes *et al.* studied the solution chemistry of 244 copper arsenate mineral [9]. These authors reported the relationship between 245 euchroite Cu₂(AsO₄)(OH)³H₂O, strashimirite Cu₈(AsO₄)₄(OH)₄⁻4H₂O and arhbarite 246 Cu₂Mg(AsO₄)(OH)₃. Complex paragenetic relationships exist between these copper 247 arsenate minerals. An example of this relationship exists for the minerals of the 248 Zapachitsa deposit in Bulgaria. The sequence of deposition of the copper arsenate 249 minerals is proposed to be cornwallite-olivenite-tyrolite-strashimirite-conichalcite. It 250 is possible that the mineral sample reported based upon infrared spectra, by Stoilova 251 and Minceva-Stefanova may be likely a mixture of some minerals from this mineral 252 paragenetic sequence.

253

254 The Raman spectra of strashimirite A and B in the 100 to 600 cm⁻¹ region are 255 shown in Fig. 5. A broad band centred upon 500 cm^{-1} for strashimirite A, which may 256 be resolved into component bands at 467 , 497 , 526 and 554 cm⁻¹ probably includes 257 bands attributable to the v_4 (AsO₄³) bending mode. The observation of multiple 258 bands suggests that the symmetry of the $(AsO₄³)$ ion has been reduced. The shoulder 259 band at 467 cm⁻¹ is probably ascribed to the CuO stretching mode. The higher

- 260 wavenumber band may be assigned to a water librational mode. A band in this
- 261 position is expected although in the Raman spectrum may be of low intensity. A
- 262 Raman band is observed at 560 cm⁻¹ for strashimirite B. The band at 507 cm⁻¹ for
- 263 strashimirite B is broad and is assigned to the $v_4 (AsO_4)^3$ bending mode. Stoilova and

264 Minceva-Stefanova reported infrared bands at 430, 470 and 92 and assigned these 265 bands to v_4 (AsO₄³) bending mode [1]. The appearance of multiple bands in the v_4 266 (AsO₄³⁻) bending region further provides evidence for the reduction of the symmetry 267 of the $(AsO₄³)$ anion to C_{2v} or even lower.

268

269 The low wavenumber region of the basic copper arsenates is complex and it is 270 difficult to separate the Raman bands according to their symmetry. The spectrum of 271 the low wavenumber region of the basic copper arsenates may be divided into three 272 separate regions: (a) 450 to 650 cm⁻¹ (b) 400 to 450 cm⁻¹ and (c) 400 to 250 cm⁻¹. It is 273 proposed that these three regions define the (a) v_4 modes (b) v_2 modes and (c) $(AsO₄³)$ 274 lattice modes. The v_2 bending vibration should be common to all the basic copper 275 arsenates spectra and should be intense. The $(AsO₄³)$ bending mode ($v₂$) should be 276 found at around 342 cm^{-1} . In the Raman spectrum of A, two bands at 337 and 393 277 cm⁻¹ may be ascribed to this vibrational mode. Three Raman bands are observed for 278 B at 343, 374 and 411 cm⁻¹. The first band is attributed to the $v_2 (AsO₄³)$ bending 279 mode. The band at 350 cm⁻¹ for olivenite, 340 cm⁻¹ for cornwallite and at 380 cm⁻¹ is 280 assigned to the v_2 (AsO₄³) symmetric bending vibration. The observation of this 281 mode below 400 cm⁻¹ has been predicted but not reported [31]. Our results differ 282 from those of Griffith who reported bands at 324 and 310 cm⁻¹ for the $(AsO₄³)$ 283 bending modes of olivenite [33].

284

285 In the low wavenumber region of the basic copper arsenates, intense Raman 286 bands are observed (Fig. 5). In the Raman spectrum of the low wavenumber region 287 for A, Raman bands are observed at 152, 172, 220, 239 and 294 cm^{-1} . In the Raman 288 spectrum of the low wavenumber region bands for B are observed at 144, 183, 219, 289 242 and 299 cm⁻¹. The intense sharp band at 299 cm⁻¹ is assigned to the AsO 290 stretching vibration. In the structure of strashimirite it is proposed (without the single 291 crystal X-ray evidence) that the $(AsO₄³)$ units are not equivalent and thus more than 292 one CuO vibration will exist. Stoilova and Minceva-Stefanova tabled results for the 293 CuO stretching vibration at 492 and 540 cm⁻¹ in the infrared spectrum [1]. For

294 olivenite intense bands are observed at 310 and 286 cm⁻¹ and are assigned to the CuO 295 stretching vibration. A second low intensity band is observed at 258 cm^{-1} and may be 296 also assigned to a CuO stretching vibration. Two CuO vibrations exist because there 297 are two types of units namely OH and AsO4 units coordinating to the copper. The 298 intense bands for A at 144 and 183 cm⁻¹ may be assigned to OCuO bending 299 vibrations. Intense Raman bands are also observed at 219 and 151 cm⁻¹ for olivenite. 198 and 172 cm⁻¹ for cornwallite and at 189 and 185 cm⁻¹ for clinoclase. These bands 301 may also be assigned to OCuO bending vibrations.

302

303 The Raman spectrum in the OH stretching region of strashimirite A and B are 304 shown in Fig. 6. The infrared spectra of the strashimirite minerals are shown in Fig. 7. The Raman spectrum of A shows three resolved bands at 3450, 3588 and 3585 cm- 305 1306 ¹. The first two bands are attributed to water stretching vibrations and the band at 307 3585 cm⁻¹ to the OH stretching vibrations of the hydroxyl units. The infrared 308 spectrum of A shows a complex set of overlapping bands. Band component analysis 309 resolves component bands at 2728, 3012, 3235 cm⁻¹ ascribed to water stretching 310 bands; and component bands at 3361 and 3514 cm⁻¹ attributed to the OH stretching 311 vibrations of hydroxyl ions. The Raman spectrum in this spectral region for B shows 312 a broad band centred on 3444 cm⁻¹ with a sharp intense band at 3497 cm⁻¹. One likely 313 assignment is that the first band is due to the water stretching vibration and the latter 314 to the stretching vibration of the hydroxyl units. As for strashimirite A, a complex set 315 of overlapping infrared bands are observed. Infrared bands are observed at 2731, 316 3059 and 3273 cm⁻¹ assigned to water stretching bands and at 3439 and 3579 cm⁻¹ 317 assigned to the OH unit stretching vibrations of the hydroxyl ions.

318

319 Stoilova and Minceva-Stefanova reported the infrared OH stretching bands at 320 3380, 3420 and 3500 cm^{-1} [1]. These values differ from those reported in Fig. 5. 321 These authors considered that these bands provided evidence of strong hydrogen 322 bonds in the strashimirite mineral. Low intensity infrared bands found by these 323 authors were observed at 2133, 2740 and 2900 cm^{-1} and provided further evidence of 324 strong hydrogen bonds in the strashimirite structure. In our Raman spectrum of 325 strashimirite low intensity bands were found at 2124, 2346, 2924 and 3058 cm⁻¹. In 326 our infrared spectrum, bands were observed at around 2730 and 2842 cm⁻¹ supporting 327 the proposition by Minceva-Stefanova for strong hydrogen bonding in the

328 strashimirite structure. The observation of these bands in our Raman and infrared 329 spectra supports the concept of strong hydrogen bond interactions between the water 330 units and the arsenate anions.

331

332 A comparison may be made with the Raman spectrum of the hydroxyl 333 stretching region of other basic copper arsenates (Table 3). The Raman spectrum of 334 olivenite displays two overlapping bands centred upon 3464 and 3437 cm⁻¹; and for 335 cornwallite two partially band-separated peaks are observed at 3411 and 3350 cm⁻¹. 336 In the Raman spectrum of cornubite, two bands are observed at 3324 and 3042 cm^{-1} . 337 For clinoclase two well separated bands separated peaks are observed at 3559 and 338 3339 cm^{-1} . The observation of two hydroxyl-stretching vibrations means that there 339 are two distinct and different hydroxyl units in the basic copper arsenate minerals. 340 The difference between the peak positions of olivenite and cornwallite is related to the 341 strength of the hydrogen bond formed between the hydroxyl unit and adjacent 342 arsenate units. This bonding is much stronger in cornwallite as indicated by the lower 343 wavenumber position of the hydroxyl-stretching vibrations. One interpretation is that 344 the higher wavenumber vibration is ascribed to the As-OH vibration and the lower 345 wavenumber hydroxyl stretching frequency to the As-OH[.]O vibration.

346

347 The infrared spectra of the strashimirite minerals A and B are shown in Fig. 8. 348 The 1300 to 1800 cm⁻¹ spectral profile for A, may be resolved into component bands 349 at 1617, 1649 and 1673 cm^{-1} with additional low intensity bands at 1420 and 1459 350 cm⁻¹. These latter two bands may indicate the presence of carbonate anions. 351 However, no carbonate ions were inferred from the analysis. It is likely these low 352 intensity infrared bands are due to overtone or combination bands. The infrared band 353 at 1617 cm⁻¹ is typical of weakly hydrogen bonded water and the two bands at 1649 354 and 1673 cm⁻¹ are attributed to strongly hydrogen bonded water. The identification of 355 these bands supports the concept that water is strongly bonded in the structure of 356 strashimirite and is required for the stability of the mineral. The infrared spectrum of 357 strashimirite B also shows bands at 1624 and 1674 cm^{-1} , which supports the concept 358 of two types of water in the strashimirite structure.

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361 *3.2. Diagenetic relationships between the basic copper arsenate minerals*

363 Williams [10] has presented phase stability relationships between a range of

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- 458
- 459
- 460

461 **Table 1**

462 **Refined unit-cell parameters for strashimirite**

Occurrence		a/A	$b/\text{\AA}$	c/A	β / ^o	Volume/ A^3
Zálesí (CR)		9.56(1)	18.38(3)	9.10(1)	$97.26(9)^{\circ}$	1587(4)
Lubietová (SR)		9.524(3)	$18.536(6)$ 9.058(4)		96.96(4)	1587.1(9)
Zapačica	[35]	9.70	18.90	9 1 2 7	97 25	1660
Novoveská Huta		$[28, 36]$ 9.708(6) 18.84(9)		8.936(7)	97.22(5)	16214
Redziny	$[37]$	9.719(2)	$18.806(5)$ 8.937(3)		97.31(3)	1620.2

463

465 **Table 2**

		Lubietová (SR) ^a		Zálesí $(CR)^{b}$			
	mean	range	mean	range			
CaO	0.06	$0.04 -$ 0.07	0.51	$0.38 -$	0.65	0.00	
FeO	0.00	$0.00 -$ 0.00	0.06	0.00 \sim	0.13	0.00	
CuO	50.05	$50.03 - 50.08$	50.31	$47.80 -$	52.57	52.07	
CoO	0.02	$0.02 -$ 0.02	0.30	$0.26 -$	0.38	0.00	
N _i O	1.72	$1.72 -$ 1.73	0.54	$0.42 -$	0.69	0.00	
ZnO	0.22	0.27 $0.16 -$	0.38	$0.31 -$	0.43	0.00	
Al_2O_3	0.17	0.19 $0.16 -$	0.06	$0.00 -$	0.12	0.00	
Sb_2O_3	0.02	$0.02 -$ 0.03	0.08	0.05 \sim	0.11	0.00	
SiO ₂	0.09	$0.07 -$ 0.11	0.09	$0.03 -$	0.21	0.00	
As ₂ O ₅	35.53	$35.27 - 35.80$	37.20	36.22 \overline{a}	38.26	37.61	
P_2O_5	0.88	$0.86 -$ 0.90	0.04	$0.02 -$	0.05	0.00	
SO ₃	0.17	0.18 $0.17 -$	0.55	$0.45 -$	0.64	0.00	
H_2O^*	10.46	$0.00 -$ 0.00	10.44			10.32	
total	99.39		100.57			100.00	

466 **Chemical composition of strashimirite (wt. %)**

^a 467 mean and range of two point analyses

 468 b mean and range of six point analyses

469 composition for ideal formula $Cu_8(AsO_4)_4(OH)_4$:5H₂O

 470 H₂O^{*} - content calculated on the basis of charge balance and theoretical content 5

471 H2O.

474 **Table 3 Results of the vibrational spectroscopic analysis of strashimirite and in**

475 **comparison with the Raman spectra of other basic copper arsenates**

