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The Largest Volcanic Eruptions on Earth

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1 Abstract

2
3 Large igneous provinces (LIPs) are sites of the most frequently recurring, largest volume basaltic
4 and silicic eruptions in Earth history. These large-volume ($>1,000 \text{ km}^3$ dense rock equivalent)
5 and large-magnitude ($>M8$) eruptions produce areally extensive (10^4 - 10^5 km^2) basaltic lava flow
6 fields and silicic ignimbrites that are the main building blocks of LIPs. Available information on
7 the largest eruptive units are primarily from the Columbia River and Deccan provinces for the
8 dimensions of flood basalt eruptions, and the Paraná-Etendeka and Afro-Arabian provinces for
9 the silicic ignimbrite eruptions. In addition, three large-volume (675 - $2,000 \text{ km}^3$) silicic lava flows
10 have also been mapped out in the Proterozoic Gawler Range province (Australia), an interpreted
11 LIP remnant. Magma volumes of $>1,000 \text{ km}^3$ have also been emplaced as high-level basaltic and
12 rhyolitic sills in LIPs. The data sets indicate comparable eruption magnitudes between the
13 basaltic and silicic eruptions, but due to considerable volumes residing as co-ignimbrite ash
14 deposits, the current volume constraints for the silicic ignimbrite eruptions may be considerably
15 underestimated. Magma composition thus appears to be no barrier to the volume of magma
16 emitted during an individual eruption. Despite this general similarity in magnitude, flood basaltic
17 and silicic eruptions are very different in terms of eruption style, duration, intensity, vent
18 configuration, and emplacement style. Flood basaltic eruptions are dominantly effusive and
19 Hawaiian-Strombolian in style, with magma discharge rates of $\sim 10^6$ - 10^8 kg s^{-1} and eruption
20 durations estimated at years to tens of years that emplace dominantly compound pahoehoe lava
21 flow fields. Effusive and fissural eruptions have also emplaced some large-volume silicic lavas,
22 but discharge rates are unknown, and may be up to an order of magnitude greater than those of
23 flood basalt lava eruptions for emplacement to be on realistic time scales (<10 years). Most
24 silicic eruptions, however, are moderately to highly explosive, producing co-current pyroclastic
25 fountains (rarely Plinian) with discharge rates of 10^9 - $10^{11} \text{ kg s}^{-1}$ that emplace welded to
26 rheomorphic ignimbrites. At present, durations for the large-magnitude silicic eruptions are
27 unconstrained; at discharge rates of 10^9 kg s^{-1} , equivalent to the peak of the 1991 Mt Pinatubo
28 eruption, the largest silicic eruptions would take many months to evacuate $>5,000 \text{ km}^3$ of magma.
29 The generally simple deposit structure is more suggestive of short-duration (hours to days) and
30 high intensity ($\sim 10^{11} \text{ kg s}^{-1}$) eruptions, perhaps with hiatuses in some cases. These extreme
31 discharge rates would be facilitated by multiple point, fissure and/or ring fracture venting of

1 magma. Eruption frequencies are much elevated for large-magnitude eruptions of both magma
2 types during LIP-forming episodes. However, in basalt-dominated provinces (continental and
3 ocean basin flood basalt provinces, oceanic plateaus, volcanic rifted margins), large magnitude
4 (>M8) basaltic eruptions have much shorter recurrence intervals of 10^3 - 10^4 years, whereas similar
5 magnitude silicic eruptions may have recurrence intervals of up to 10^5 years. The Paraná-
6 Etendeka province was the site of at least nine >M8 silicic eruptions over an ~1 Myr period at
7 ~132 Ma; a similar eruption frequency, although with a fewer number of silicic eruptions is also
8 observed for the Afro-Arabian Province. The huge volumes of basaltic and silicic magma erupted
9 in quick succession during LIP events raises several unresolved issues in terms of locus of
10 magma generation and storage (if any) in the crust prior to eruption, and paths and rates of ascent
11 from magma reservoirs to the surface.

12
13 Available data indicate four end-member magma petrogenetic pathways in LIPs: 1) flood basalt
14 magmas with primitive, mantle-dominated geochemical signatures (often high-Ti basalt magma
15 types) that were either transferred directly from melting regions in the upper mantle to fissure
16 vents at surface, or resided temporarily in reservoirs in the upper mantle or in mafic underplate
17 thereby preventing extensive crustal contamination or crystallisation; 2) flood basalt magmas
18 (often low-Ti types) that have undergone storage at lower \pm upper crustal depths resulting in
19 crustal assimilation, crystallisation, and degassing; 3) generation of high-temperature anhydrous,
20 crystal-poor silicic magmas (e.g., Paraná-Etendeka quartz latites) by large-scale AFC processes
21 involving lower crustal granulite melting and/or basaltic underplate remelting; and 4)
22 rejuvenation of upper-crustal batholiths (mainly near-solidus crystal mush) by shallow intrusion
23 and underplating by mafic magma providing thermal and volatile input, producing large volumes
24 of crystal-rich (30-50%) dacitic to rhyolitic magma and for ignimbrite-producing eruptions, well-
25 defined calderas up to 80 km diameter (e.g., Fish Canyon Tuff model), and characteristic of some
26 silicic eruptions in silicic LIPs.

27
28 **Keywords: supereruption; Large Igneous Provinces; flood basalt; rhyolite; ignimbrite**

29 30 **1. Introduction**

31
32 The generation and emplacement of large igneous provinces (LIPs) are anomalous transient

1 igneous events in Earth's history resulting in rapid and large volume accumulations of volcanic
2 and intrusive igneous rock (Coffin & Eldholm, 1994; Bryan & Ernst, 2008). LIP events have
3 been estimated to have had a frequency of one every 20 Myrs since the Archean (Ernst &
4 Buchan, 2002), but when the current oceanic LIP record dating back to 250 Ma is also included,
5 this frequency is reduced to one per 10 Myr (Coffin & Eldholm, 2001; see also Prokoph & Ernst,
6 2004). The volcanic and intrusive products of individual LIPs collectively cover areas well in
7 excess of 0.1 Mkm², and typically, extruded volcanic deposit volumes are ≥ 1 Mkm³. Oceanic
8 plateaus define the upper limits of the areal and volumetric dimensions of terrestrial LIPs, with
9 reconstruction of the Ontong-Java, Hikurangi and Manihiki plateaus (Taylor, 2006) having a pre-
10 rift areal extent of ~ 3.5 Mkm², larger than the Indian sub-continent, a maximum crustal thickness
11 of 30 km and a total possible maximum igneous volume of 59-77 Mkm³ (Kerr & Mahoney,
12 2007).

13
14 A distinguishing feature of LIPs, as exemplified by continental flood basalt provinces, is the high
15 magma emplacement rates (e.g., Storey et al., 2007) where aggregate magma volumes of ≥ 1
16 Mkm³ are emplaced from a focussed source during 1 to 5 million year-long periods or pulses
17 (Bryan & Ernst, 2008). In detail, most LIP eruptions had magnitudes significantly greater than
18 those of historic eruptions, tending towards extraordinarily large-volume eruptions ($>10^3$ km³;
19 Tolan et al., 1989; Jerram, 2002; White et al., 2009; Chenet et al., 2009), making LIP volcanism
20 exceptional. However, to produce such tremendous cumulative volumes of erupted magma, the
21 short-lived, main eruptive pulses of LIP events must consist of many and frequently recurring,
22 large-volume eruptions, each evacuating 10^2 - 10^3 km³ of magma. Consequently, it is the volume
23 of magma emitted during these individual eruptions, the frequency of such large-volume
24 eruptions, and the total volume of magma intruded and released during the main igneous pulses
25 that make LIP events so exceptional in Earth history, and called upon to explain environmental
26 and climatic changes and mass extinctions (e.g., Rampino & Stothers, 1998; Courtillot, 1999;
27 Courtillot & Renne, 2003; Wignall, 2001; 2005; Self et al., 2005; Kelley, 2007).

28
29 Despite the total cumulative erupted volumes and timing of LIP events being reasonably well-
30 constrained (Coffin & Eldholm, 1994; Bryan & Ernst, 2008), our current understanding of the
31 size, duration and frequency of individual LIP eruptions is very limited. Considerable focus has
32 been on flood basalt eruptions in the continental flood basalt provinces, which are the best

1 exposed and studied examples of LIPs. Almost all information on the size of individual flood
2 basaltic eruptions comes from the many studies undertaken on the Columbia River Flood Basalt
3 Province, which is the smallest ($\sim 0.234 \text{ Mkm}^3$) and youngest example of a continental flood
4 basalt province (e.g., Swanson et al., 1975; Reidel et al., 1989; Tolan et al., 1989; Self et al.,
5 1997; Camp et al., 2003; Hooper et al., 2007). It is only recently that some understanding has
6 been made on the magnitude of flood basalt eruptions from other flood basalt provinces (Deccan:
7 Jay & Widdowson, 2006; Self et al., 2008; Chenet et al., 2009).

8
9 By contrast, similarly large-volume silicic volcanic eruptions are known from a number of
10 tectonic regimes, but which are exclusively continental in crustal setting. Extension of active
11 continental margins, whether in narrow, rifted arc or back-arc settings (e.g., Taupo Volcanic
12 Zone) or broader extensional belts (e.g., Basin & Range Province, western USA) and intraplate
13 to rifted continental environments (e.g., Afro-Arabian province) have been the most productive
14 settings for large-volume ($>1,000 \text{ km}^3$) silicic eruptions since the middle Tertiary (Mason et al.,
15 2004). Consequently, unlike flood basalt eruptions, large-volume silicic eruptions are not
16 exclusive to LIPs and not restricted to discrete eruptive episodes such as LIP events throughout
17 Earth history (Thordarson et al., 2009). The presently determined average recurrence rate of one
18 silicic eruption of Magnitude 8 or greater (Pyle, 1995; 2000) every 100,000-200,000 years (Self,
19 2006) reflects the contribution from sources in a variety of tectonic settings. This relatively
20 higher frequency for large magnitude silicic eruptions means that they pose a greater hazard to
21 human civilization than flood basaltic eruptions (Thordarson et al., 2009). What is distinctive
22 regarding large-volume silicic eruptions from LIPs is their association with large-magnitude
23 basaltic eruptions, their enhanced frequency and the cumulative volume of silicic magma
24 emplaced (up to 10 Mkm^3) when compared to other tectonic settings (Bryan et al., 2002; Mason
25 et al., 2004; Bryan & Ernst, 2008).

26
27 Super-eruptions have recently been defined as those yielding more than $1 \times 10^{15} \text{ kg}$ of magma
28 (Sparks et al., 2005; Self, 2006). For rhyolitic eruptions, this is equivalent to $\sim 410 \text{ km}^3$ (at a
29 magma density of $2,450 \text{ kg m}^{-3}$). However, super-eruption has not yet been strictly applied to
30 basaltic eruptions, where only 360 km^3 of erupted magma is required, given their higher magma

1 density of $\sim 2,750 \text{ kg m}^{-3}$. Here we present a compilation of known eruption volumes for the very
2 largest ($>$ Magnitude 8.5, equivalent to $\sim 1,160 \text{ km}^3$ of basalt lava, or $\sim 1,280 \text{ km}^3$ of dense silicic
3 lava or ignimbrite) basaltic and rhyolitic eruptions from LIPs. This is to complement recent
4 compilations for example, on the largest Tertiary-Quaternary silicic explosive eruptions (Mason
5 et al., 2004), and provide a basis for improved long-term eruption rate estimates (e.g., White et
6 al., 2006), which are currently based on sparse data from LIPs. Consequently, our understanding
7 of what are the largest eruptions is limited and biased to late Tertiary and Recent volcanic
8 activity. In these recent compilations, only one silicic eruption of magnitude 9 ($1 \times 10^{16} \text{ kg}$, or
9 $\sim 5,000 \text{ km}^3$ of dense magma) has been recognized, which occurred 28 Ma (Lipman et al., 1970;
10 Mason et al., 2004). Issues investigated with the data set presented here are: What are the largest
11 eruptions? What are the physical limits that may exist for eruption magnitude and whether
12 magma composition imposes any limitation? How do basaltic and silicic “super-eruptions” differ
13 in terms of eruption mechanisms, rates, durations and frequencies during LIP-forming events?
14 And what implications do these issues have for the generation and storage of such prodigious
15 magma volumes?

16 17 **2. Determining the products of single eruptions**

18
19 Determining what deposits constitute the products of an individual eruption and assessing
20 erupted volumes are not straightforward in LIPs given exposure problems (e.g., concealment,
21 burial or uplift and erosion), the potential great extent of (often thin) eruptive units (10^4 - 10^6 km^2),
22 tectonic deformation and fragmentation, and subtle lithologic or geochemical distinction.
23 Important tools for discriminating individual eruptive units are superposition, the presence or
24 absence of internal palaeosoils, sedimentary or other lithologically distinct deposits (Figs 1, 2),
25 petrography, mineral chemistry, bulk rock or juvenile component compositional characteristics,
26 and paleomagnetic character (see Milner et al, 1995; Self et al., 1997; Jay et al., 2009; Chenet
27 et al., 2009). Geochemistry has been useful at regional scales in defining stratigraphic units of
28 basaltic lavas with distinctive geochemical features (e.g., Mangan et al., 1986; Devey &
29 Lightfoot, 1986; Peate et al., 1992; Hooper, 1997; Marsh et al., 2001) but greatest success is
30 achieved where a combination of approaches are used (e.g., Chenet et al., 2008; 2009; Jay et al.,

1 2009). Examples of significant regional correlations include the dissected and tectonically
2 fragmented silicic eruptive units of the Paraná-Etendeka flood basalt province (Milner & Duncan,
3 1987; Milner et al., 1992; 1995) and on-land ignimbrites of the Afro-Arabian province with distal
4 co-ignimbrite ash deposited in marine basins (Ukstins Peate et al., 2003; 2008).

5
6 A complication in unravelling the products of an individual eruption in LIPs is caused by factors
7 such as the great areal extent of volcanism, multiple vent activity, potentially long crustal
8 transport distances, and in particular, the duration of flood basalt eruptions. Given these factors,
9 synchronicity of contrasting magmas (Fig. 3) and of widely separated eruptions are a distinct
10 possibility. Because of the high frequency of large eruptions during a LIP-forming event, the
11 products of the majority of eruptions cannot be discriminated on the basis of current radiometric
12 dating techniques. Heterogeneous granitoids produced by mingled gabbroic and granitic magmas
13 in associated intrusive complexes (e.g., Messum; Fig. 3A) record the simultaneous existence of
14 mafic and silicic magmas in LIP plumbing systems (Vogel, 1982; Ewart et al., 2002).
15 Consequently, it remains unclear where interbedded units do exist whether this interstratification
16 implies either a substantial timebreak or simultaneous (basalt-rhyolite or basalt-basalt) eruptions
17 from different vent regions within the province.

18
19 Two pertinent examples of the above occur in the Paraná-Etendeka flood basalt province (Fig.
20 4): 1) two basaltic magma types (Urubici and Gramado) are interbedded in the southern Paraná
21 (Peate et al., 1999); and 2) two discontinuous and chemically different basalt lavas are
22 interbedded within the Goboboseb Quartz Latite of the Etendeka Province (Milner & Ewart,
23 1989; Milner et al., 1992; 1995). In the Paraná example, compositionally identical Urubici-type
24 basaltic lava flow units 9 and 10, exposed along the coastal escarpment, were divided into two
25 units as they are separated by a Gramado-type basaltic lava flow near Urubici. We suggest that
26 units 9 & 10 were part of a single eruptive event (sourced from a vent region to the north where
27 Urubici-type magmas predominate) that coincided with the simultaneous eruption of a Gramado
28 flow field from the south.

29
30 For flood basalt eruptions, the term *lava flow field* refers to the aggregate product of a single

1 eruption and can be formed of one or more lava flows, each the product of a vent or a group of
2 vents along a fissure segment (Self et al., 1997). Consequently, all flood basalt lava fields are
3 compound pahoehoe lavas composed of 1000's of lava sheet lobes, and innumerable smaller
4 lobes, stacked/superposed in some places and laterally arranged in others. By contrast, the
5 stratigraphy of silicic eruptive units appears much simpler, as exemplified by the Paraná-
6 Etendeka quartz latites (Fig. 5), which crop out as relatively featureless extensive sheets with a
7 simple internal structure and jointing, leading to the interpretation as single flow or cooling units
8 (Milner et al., 1992). A simple deposit structure, an extremely chemically uniform or distinctive
9 character and chemical distinction from other silicic units are used here to support the
10 interpretation of silicic units in flood basalt provinces such as the Paraná-Etendeka, Karoo and
11 Afro-Arabian provinces as being the products of individual eruptions.

12 13 *2.1 Volume estimations*

14
15 Volume estimation of LIP eruptive units is a difficult task, and is strongly affected by erosion,
16 burial and concealment of eruptive units, and tectonic fragmentation of the provinces. It requires
17 excellent exposure, little deformation or erosion of the sequences, and textural and/or
18 geochemical distinction to aid in correlation. As a consequence of these factors, each LIP will
19 have a different prospectivity potential for delineating the products of large-magnitude eruptions
20 (Table 1). For the basaltic eruptive units, long-distance crustal transport in sill-dyke complexes
21 and multiple venting sites (e.g., Scarab Peak lavas, Ferrar LIP, Elliot et al., 1999) complicate this
22 task. Additional complications arise for the silicic eruptive units in that nonwelded pyroclastic
23 material is more easily eroded than massive or even autobrecciated lava, and that deposit
24 volumes may reside as intracaldera, outflow, tephra fallout (from Plinian and/or co-ignimbrite
25 ash plumes) deposits (Mason et al., 2004; Self, 2006; de Silva, 2009), as voluminous
26 resedimented pyroclastic deposits (Bryan et al., 1997), and substantial volumes may actually
27 reside as tephra deposits in ocean basins (e.g., Straub & Schmincke, 1998; Pattan et al., 2002;
28 Ukstins Peate et al., 2005).

29
30 A large proportion of the erupted volume may reside in co-ignimbrite ash deposits deposited
31 >1,000 km downwind (Sparks & Walker, 1977; Moore, 1991; Self, 1992; Koyaguchi & Tokuno,

1993). Co-ignimbrite ash plumes are generated where the pyroclastic suspension currents lift off at their run-out distance (Freundt, 1999), which may be ≥ 100 km from source, in the largest LIP silicic eruptions. Modelling has predicted that a substantial fraction (30-80%) of the erupted mass is entrained in to co-ignimbrite ash plumes that reach stratospheric heights where the erupted material is then distributed widely (Woods & Wohletz, 1991; Bonnadonna et al., 1998). This occurs even during eruptions of relatively low magmatic gas content and high discharge (Freundt, 1999), which may be more characteristic of silicic eruptions in flood basalt provinces. Recent insight into this issue, and the potential extent and increased magnitude of silicic eruptions from LIPs comes from the correlation of deep sea ashes with onshore ignimbrites from the Afro-Arabian LIP (Ukstins Peate et al., 2003; 2008). Conversely, approximately 200 very widely dispersed basaltic to rhyolitic ash layers are preserved in basins adjacent to the North Atlantic LIP (Larsen et al., 2003), but source regions and total erupted volumes remain very poorly constrained.

Deposit dense-rock-equivalent volumes from known large-volume LIP eruptions are listed in Tables 2 & 3. These should be considered as minima. Although studies of some recent >M8 caldera-forming explosive eruptions have indicated that deposit volumes reside in approximately equal proportions as intracaldera (I), outflow (O) and co-ignimbrite ash (A) deposits (Rose & Chesner, 1987; Wilson, 2001), we have not adjusted deposit volumes (applying the assumption that total volume = I+O+A) as suggested by Mason et al. (2004). This is because for the silicic eruptions listed in Table 3: 1) calderas and any potential ignimbrite fill have rarely been identified in LIPs; 2) it remains unclear if these LIP eruptions did lead to true caldera collapse (and what the relative timing of collapse was) and eruptions may have instead produced large regional downsagging (e.g., Ewart et al., 2002; Mahoney et al. 2008); and 3) while recent efforts have identified distal co-ignimbrite ash deposits for some Afro-Arabian silicic eruptions (Ukstins Peate et al., 2003; 2008) allowing an upward revision of eruptive volumes on the basis of O+A (Table 3), distal co-ignimbrite ash deposits have yet to be positively identified for the Paraná-Etendeka silicic eruptive units. Silicic ash horizons present in some drill cores recovered from the South Atlantic have been dated to an approximate age consistent with Paraná-Etendeka eruptions (Barker et al., 1988). These offer hope that distal ash deposits do exist but substantiation requires further detailed analysis and geochemical comparison (Mawby, 2008).

1 It therefore should be noted that deposit volumes listed in Table 3 for the Paraná-Etendeka silicic
2 eruptive units equate to approximately two-thirds of the total erupted volume (O+I) if co-
3 ignimbrite ash deposits exist.

4 5 **3. LIP eruptions**

6
7 LIPs are dominantly basaltic igneous events and the primary building blocks are extensive (10^3 -
8 10^5 km²), sheet-like lava flow fields or sills (Fig. 6; e.g., Self et al., 1997; Jerram, 2002; White
9 et al., 2009). However, in continental LIPs (flood basalt provinces and volcanic rifted margins),
10 mafic volcanoclastic deposits (Ross et al., 2005; Ukstins Peate & Bryan, 2008), sill and dyke
11 intrusions (Jerram, 2002; Elliot et al., 2008) and silicic ignimbrites (Bryan et al., 2002; Bryan,
12 2007) are also, areally and volumetrically significant and important architectural components
13 (White et al., 2009). The flood basalts and rhyolites of LIP events have remarkably similar
14 volumes for individual eruptions (Tables 2, 3).

15 16 *3.1 Flood basalt eruptions in LIPs*

17
18 The basaltic lava flow fields in LIPs in many cases are so extensive, and the provinces so
19 widespread and fragmented or eroded, that it has taken many years of study to determine what
20 the products of a "typical" flood basalt eruption comprise (White et al., 2009). Most insight into
21 the products of flood basalt eruptions has come from studies on the Columbia River Basalt
22 province (CRB) of northwest USA (see summaries in Tolan et al., 1989; Reidel et al., 1989; Self
23 et al., 1997; Hooper, 1997).

24
25 Dominantly pahoehoe lava flow fields form the flood basalt lavas in LIPs. Each flow field is
26 interpreted to be the product of one, but potentially sustained (Self et al., 1997), eruptive event
27 producing several major lava flows that in turn consist of multiple sheet-like flow lobes. The
28 major sheet lobes contain the majority of the lava volume, and in the CRB are commonly 20 to
29 30 m thick, several kilometres wide. In the Karoo and Etendeka flood basalt provinces, however,
30 studied sections generally reveal only a few lava flows >20 m thick, suggesting that the majority
31 of the flood lava volume is not always expressed in thick lava flow units. Nevertheless, the flood

1 basalts from these provinces show features consistent with *in situ* flow thickening by endogenous
2 growth and inflation (Self et al., 1997; Bondre et al. 2004; White et al., 2009). These extensive
3 lobes, with aspect ratios (length/thickness) ranging from ~50 to 500, are the basic building-blocks
4 of continental flood basalt provinces and give the provinces their "layer-cake", or step-like,
5 appearance (Fig. 6A; Jerram & Widdowson, 2005; White et al., 2009). Sheet lobes show a
6 relatively simple internal structure of crust and lava core facies (Fig. 6B, C) that are ubiquitous
7 in the lavas regardless of thickness, and persist over the extent of the entire flow field (10^2 - 10^3
8 km^2 ; Self et al., 1997; 2008). Parts of some continental flood basalt provinces, e.g., Deccan, are
9 formed of flow-fields dominated by thinner, smaller lava units stacked up to form compound
10 lavas (Walker, 1972; Bondre et al., 2004). Complexities can occur where lava enters lakes or
11 oceans producing pillowed, hyaloclastite and pyroclastic facies (e.g., Waters, 1960; Larsen et al.,
12 2006).

13
14 Volcanological studies in the CRB province on the eruptive characteristics of flood basalt
15 eruptions have shown that: 1) at least parts of flood basalt eruptions were Hawaiian-like in nature
16 at the vent (Reidel & Tolan, 1992); 2) the flow fields most likely originated from prolonged
17 eruptions lasting years to decades (Self et al., 1996, 1997, 1998; Thordarson & Self, 1996, 1998),
18 with the larger flow fields fed by very long fissures (e.g., Swanson et al., 1975); 3) the length of
19 flow fields has been supply-limited rather than landscape- or cooling-limited (Stephenson et al.,
20 1998; Keszthelyi & Self, 1998; Keszthelyi et al., 2004; Self et al., 2008), with several reaching
21 the sea >400 km from vent; 4) at any one time, eruptive activity was more likely confined to
22 distinct fissure segments on the vent system (Thordarson & Self, 1998), resulting in incremental
23 flow advancement and flow field growth; and 5) postulated eruption rates of $\sim 4,000 \text{ m}^3 \text{ s}^{-1}$ (Self
24 et al., 1998) are similar to the maximum sustained eruption rates of the largest historic eruptions
25 (e.g., Laki, Thordarson & Self, 1993). Given the inferred duration and sporadic and restricted
26 activity along the vent fissure system of flood basalt eruptions, defining what the product of one
27 eruption is remains a difficult task in flood basalt studies.

28
29 Flood basalt lavas vary from being aphyric or sparsely phyric ($\leq 5\%$) to strongly porphyritic (up
30 to 40%) with plagioclase and clinopyroxene usually dominating the phenocryst assemblage.
31 Generally, geochemical data define two compositionally distinct groups of mafic lavas that can

1 be clearly differentiated in terms of immobile high field strength elements (Zr, Ti, Y, Nb) and
2 have been termed low- and high-Ti (>2 wt% TiO₂) basalts (e.g., Duncan et al., 1984; Peate et al.,
3 1992; Melluso et al., 1995; Ewart et al., 2004a). High-Ti basalts are mildly alkaline whereas the
4 low-Ti basalts are tholeiitic in composition. The presence of abundant plagioclase in the flood
5 basalts has generally been interpreted to reflect crystallisation at crustal pressures (Cox, 1980),
6 whereas the locally phenocrystic character of some high-Ti lavas has been attributed to the
7 accumulation of mafic phenocryst phases (Hooper, 1997; Ewart et al., 2004a). The aphyric lavas
8 reflect removal of crystals by fractional crystallisation, which is supported by lower MgO and
9 higher SiO₂ contents (Hooper, 1997; Ewart et al., 1998a; 2004a).

11 *3.2 Silicic eruptions in LIPs*

13 Most Quaternary examples of the products of large-volume silicic eruptions (>10 to 1,000's km³)
14 are related to Plinian and pyroclastic flow-forming explosive eruptions where a regular eruptive
15 sequence of widespread pumice and ash fallout accompanied, and was followed by, the
16 emplacement of voluminous ignimbrites leading to caldera collapse, and later, post-collapse
17 extrusion of relatively small-volume rhyolitic lava flows and domes (e.g., Valles, Taupo,
18 Yellowstone, Long Valley calderas; Walker, 1981; Miller & Wark, 2008; Wilson, 2008; cf.
19 Branney et al., 2008). Plinian fall deposits, which have accompanied many of the recent
20 examples of silicic super-eruptions (Wilson, 2008) are extremely rare in LIPs. This is despite
21 rapid accumulation rates of the volcanic piles and limited evidence for intra-LIP erosion that
22 might remove such deposits. Some examples are known of distal ash deposits preserved in
23 adjacent sedimentary or marine basins that correlate to the terrestrial flood volcanism (Heister
24 et al., 2001; Larsen et al., 2003) and in some cases are co-ignimbrite in origin (Ukstins Peate et
25 al., 2003; 2005).

27 Rhyolite lavas tend to be small volume (<100 km³) and uncommon in the continental flood basalt
28 provinces, but although more common in the silicic LIPs (Bryan et al., 2002; Bryan, 2007), are
29 similarly small in volume <100 km³. Volumetrically significant accumulations and individual
30 lava units (up to ~2,000 km³) are known from a few Proterozoic provinces (Rooiberg
31 Felsite/Bushveld igneous complex/LIP, South Africa, Twist & French, 1983; North Shore

1 Volcanic Group/Keeweenawan Rift, Minnesota, Green & Fitz, 1993), and the most detailed
2 studies have recently been published on the lavas from the Gawler Range Volcanics, South
3 Australia (Allen & McPhie, 2002; Allen et al., 2003; 2008). In addition to these eruptive
4 products, similarly extensive and voluminous high-level, lava-like rhyolitic sills have been
5 described and interpreted from ancient and partly exhumed LIPs (Trendall, 1995).

6
7 Large-volume rhyolitic ignimbrites in LIPs vary in deposit facies characteristics from extremely
8 lava-like, but showing ignimbrite depositional geometries (Fig. 5; e.g., Paraná-Etendeka quartz
9 latites, Milner et al., 1992; Mawby et al., 2006; Mawby, 2008) to “Snake River Plain”-type
10 (Branney et al., 2008) or high-grade ignimbrites with some preservation of pyroclastic textures
11 (e.g., Karoo rhyolites; Bristow, 1976; Cleverly, 1977; 1979) to more commonplace, low- to high-
12 grade welded ignimbrites containing abundant pumice lapilli and fiamme, rock fragments and
13 well-preserved vitriclastic textures (e.g., North Atlantic Igneous Province, Bell & Emeleus, 1988;
14 Afro-Arabian province, Ukstins Peate et al., 2005; Silicic LIPs, Bryan et al., 2002; Bryan, 2007).
15 This textural variation is most likely related to differing eruptive and emplacement temperatures
16 and volatile contents of the magmas (Kirstein et al., 2001; Ukstins Peate et al., 2005; Branney
17 et al., 2008). A common characteristic of the major silicic eruptive units in LIPs is that they
18 generally form relatively massive, monotonous sheet-like ignimbrite units (Fig. 5A, C), lacking
19 evidence for internal bedding or time breaks (cf. Wilson, 2008). The silicic eruptive units do
20 show lateral thickness variations due to ponding relationships in topographic (Fig. 5B) or
21 structural depressions or with proximity to vent (Ewart et al., 1998b; 2004b; Ukstins Peate et al.,
22 2005).

23
24 Additional features shared by the large-volume rhyolites in LIPs are their: 1) silica-rich bulk
25 chemical composition (~68-75 wt% SiO₂); and 2) general lateral and chemical homogeneity, yet
26 they are chemically distinct from other interbedded silicic eruptive units (e.g., Milner et al., 1995;
27 Marsh et al., 2001; Trendall, 1995; Miller & Harris, 2007). Significantly, phenocryst assemblage
28 is not diagnostic of eruptive volume, being highly variable and ranging from typically crystal-
29 poor (<10% modal phenocrysts) and anhydrous assemblages in the continental flood basalt
30 provinces to ignimbrites with high phenocryst contents (up to 40%) and hydrous mineral
31 assemblages in the silicic LIPs (Bryan, 2007), as well as those from extensional continental

1 margin settings (Mason et al., 2004). The latter ignimbrite examples are equivalent to the better-
2 known 'monotonous intermediate' ignimbrites (e.g., Hildreth, 1981; Maughan et al., 2002) such
3 as the Fish Canyon Tuff (Lipman et al., 1970; Bachmann et al., 2002). This variation in
4 phenocryst assemblage partly reflects significant differences in magma temperature with the
5 crystal-poor rhyolites from the continental flood basalt provinces being higher-temperature
6 (>1,000°C) eruptives (e.g., Milner et al., 1992; Streck & Grunder, 2008; cf. Bachmann &
7 Bergantz, 2009).

8 9 **4. Magnitude of LIP eruptions**

10
11 Constraints on flood basalt eruption magnitudes come mostly from the well-studied CRB
12 province, which is the product of many, dominantly pahoehoe flow fields varying in size from
13 1 to >2,000 km³ in volume (Tolan et al., 1989). The recent study of Reidel (2005) of proposed
14 chemically correlated flow types such as the McCoy Canyon or Cohasset (Table 2) of the Grande
15 Ronde Basalt Formation, indicates much larger volume flows may have been emplaced during
16 the interval when >60% of the volume of the CRB province was erupted (Tolan et al., 1989;
17 Camp et al., 2003). Also, studies on the Ambenali and Mahabaleshwar Formations of the Deccan
18 indicate single formations have volumes similar to the entire CRB Group (~230,000 km³, Camp
19 et al., 2003), with volumes of individual flow fields ranging from ~2,000 to >8,000 km³ (Self et
20 al., 2006; 2008). This general upper magnitude for flood basalt lavas is similar to the dimensions
21 of associated sills in flood basalt provinces, such as the enormous Peneplain Sill in the Dry
22 Valleys, Antarctica, (19,000 km²; Gunn & Warren, 1962) with an estimated volume of 4,750
23 km³, the Dufek-Forrestal intrusions (10,200-11,880 km³; Ferris et al., 1998), and the 1,500 to
24 5,000 km³ Palisades Sill of the Central Atlantic Magmatic Province (Husch, 1990; Gorrington &
25 Naslund, 1995).

26
27 The magnitude of silicic eruptions in LIP events, by contrast, has received little attention with
28 a presumption that the flood basalt eruptions were larger in volume and more likely to perturb
29 climate (cf. Cather et al., 2009). The potential scale of silicic eruptive units emplaced during LIP
30 events was first realised by the work of Milner et al. (1992, 1995) in the southern Etendeka
31 continental flood basalt province and through correlations with its rifted counterpart, the Paraná

1 province in South America. This database of extremely large silicic eruptions during LIP events
2 (Table 3) has recently been expanded following detailed studies in the northern Etendeka (Marsh
3 et al., 2001; Ewart et al., 2004b) and Afro-Arabian province (Ukstins Peate et al., 2003, 2005;
4 2008).

5
6 Areal extents of correlated quartz latite units from the Paraná-Etendeka LIP are shown in Figure
7 7. The areal extents of the quartz latite units indicated by cross-South Atlantic correlations of
8 Milner et al. (1995) and Marsh et al. (2001) are huge ($>0.1 \text{ Mkm}^2$), ranging up to 0.17 Mkm^2 and
9 greatly exceeding areas of the largest mapped flood basalt lavas from the CRB Province ($0.01\text{-}0.1$
10 Mkm^2 ; Tolan et al., 1989; Self et al., 1997). Lateral extents of the Paraná-Etendeka quartz latites
11 are also correspondingly large (up to 650 km; Milner et al., 1995; Marsh et al., 2001) and
12 equivalent to the longest run-out lengths of flood basalt lavas (Tolan et al., 1989). As emphasised
13 in previous correlations (Milner et al., 1995; Marsh et al., 2001), a distinct asymmetry exists in
14 the quartz latite distributions, occupying more area on the South American continent, as well as
15 there being a strong linear distribution to most of the quartz latite units with a preferred NW-SE
16 orientation (Fig. 7). Additionally, a spatial variation exists with the high-Ti quartz latites
17 restricted to the northern area of the Paraná-Etendeka province (Bellieni et al., 1986; Peate et al.,
18 1992; Marsh et al., 2001; Ewart et al., 2004).

19
20 Geochemical correlations for the silicic units have been revised and reassessed here since the
21 studies of Milner et al. (1995) and Marsh et al. (2001) utilising new geochemical data, and further
22 chemostratigraphic subdivisions of the Paraná rhyolites by Nardy et al. (2008). In particular, the
23 geochemical correlation of the Guarapuava and Sarusas quartz latites by Marsh et al. (2001) is
24 modified here, as there are consistent compositional differences between the major element data
25 that are most likely outside analytical error. The overall compositional similarity, however, does
26 suggest both were sourced from the same magmatic system. The recent study of Nardy et al.
27 (2008) indicates the Guarapuava, as previously defined, is a composite unit and can be
28 subdivided into two suites distinguished particularly by TiO_2 and P_2O_5 ; the Guarapuava is now
29 restricted to compositions with $\text{TiO}_2 > 1.47 \text{ wt}\%$, and the new 'Tamarana' encompasses
30 compositions with TiO_2 contents between 1.29 and 1.47 wt% (Nardy et al., 2008). Based on new
31 geochemical comparisons (Fig. 8), it is proposed here the Guarapuava correlates with the

1 Ventura, and the Tamarana with the Sarusas quartz latites from the northern Etendeka (Ewart et
2 al., 2004b). Using stratigraphic controls from the northern Etendeka province (Marsh et al., 2001;
3 Ewart et al., 2004b), these are interpreted to represent the products of different eruptions, but
4 further noting that multiple eruptive (cooling) units suggest the Sarusas quartz latite as mapped
5 by Ewart et al. (2004b), may also comprise the products of more than one eruption.

6
7 An important feature of the Paraná-Etendeka LIP is that the likely eruptive source is known for
8 two units - the Goboboseb and Springbok quartz latites, which have been correlated with the
9 PAV-A and -B rhyolites, respectively, from the Paraná (Milner et al., 1995). This has allowed
10 some perspective to be gained on run-out distances, emplacement directions, and potentially,
11 proximal-distal changes. However, as noted by Milner et al. (1995), geochemical data for the
12 Paraná PAV-A and PAV-B rhyolites, unlike other correlated silicic units, do not correspond
13 precisely with their Etendeka counterparts, and they suggested that at least for the PAV-B unit,
14 it either has been eroded from the top of the Springbok quartz latite in the Etendeka, or that it
15 represents a similar quartz latite but different eruptive unit overlying the Springbok. Ewart et al.
16 (1998b) described vertical and lateral variations in petrography for these quartz latites, and that
17 phenocryst contents decreased with distance for the Springbok PAV-B unit. Significantly,
18 variation between these Etendeka and correlative Paraná units is most pronounced for
19 phenocryst-compatible elements (e.g., TiO_2 , FeO , CaO , Sr), which show marked depletions in
20 the Paraná PAV-A and -B units; (Fig. 9) other trace elements (e.g., Rb , Th) show little to no
21 variation, and SiO_2 is also slightly enriched. These element specific variations and indications
22 for lateral variations in phenocryst content alternatively suggest that a lower degree of
23 geochemical correspondence may result from density segregation of crystals from the pyroclastic
24 density current during flow, which over a run-out distance of >300 km has resulted in detectable
25 fractionation effects in the whole-rock chemistry. Some crystal segregation layering noted by
26 Garland et al. (1995) in Paraná low-Ti (Palmas-type) silicic eruptive units may be a reflection of
27 this flow-induced density segregation. Crystal sorting during run-out has also been suggested for
28 member A of the huge M8.7 Huckleberry Ridge Tuff erupted from Yellowstone caldera at 2 Ma
29 (Christiansen, 2001), suggesting that awareness of this type of process is essential if whole-rock
30 geochemical correlations are attempted for long run-out, large-volume silicic eruptive units.

31

1 Table 3 demonstrates that dense rock equivalent eruptive volumes for many silicic eruptive units
2 in LIPs range between 1,000 and 4,000 km³ (>M8.5), and are at least equivalent to the dominant
3 erupted volumes for the largest flood basalt lavas from the CRB Province (Table 2). Importantly,
4 the data indicate the Paraná-Etendeka LIP has been the site of up to five silicic eruptions (Table
5 3) larger in magnitude than the Fish Canyon Tuff (>4,500 km³ dense rock equivalent), the most
6 commonly cited example of the largest known ignimbrite eruption (Lipman, 1997; Mason et al.,
7 2004).

8
9 In addition to large explosive silicic eruptions, recent studies of the Mesoproterozoic Gawler
10 Range Volcanic Province have delineated three >M8 silicic lava eruptions from this LIP remnant.
11 The large areal extent of these units has been aided by eruptions from multiple point or fissure
12 vents and contemporaneous eruption of chemically distinct magma batches (Allen et al., 2008;
13 McPhie et al., 2008). However, despite the similar eruption magnitudes (Table 3), these lavas
14 have run-out lengths that are over an order of magnitude less than the ignimbrite eruptive units.
15 As with basaltic sills intruded into the flood basalt piles, the 2.47 Ga Woongarra Rhyolite sill of
16 the Hammersley Basin in Western Australia (Trendall, 1995) reinforces the point that large-
17 volume batches of silicic magma are emplaced at high stratigraphic levels and available for
18 effusive or explosive eruption during LIP events. The Woongarra Rhyolite is a composite sill of
19 15,400 km³ emplaced at depths of a few hundred metres from the paleosurface. Sill emplacement
20 occurred in two pulses thought to be separated by a few hundred years, with the pulses recording
21 injection of 6,200 and 9,200 km³ of magma (Trendall, 1995).

22
23 In terms of cumulative erupted volumes, the silicic LIPs (Bryan et al., 2002; Bryan, 2007; Bryan
24 & Ernst, 2008) contain the largest volumes (~0.25-3 Mkm³) of silicic volcanic rock and are
25 equivalent to many continental flood basalt provinces. Because of this large and rapidly emplaced
26 cumulative volume, the silicic LIPs must also be host to high frequency, large volume (>1,000
27 km³ and >M8) eruptions (Mason et al., 2004; Bryan & Ernst, 2008). However, at present, there
28 are virtually no constraints on the dimensions of individual eruptions from the silicic LIPs. Mid-
29 Tertiary examples from the western USA (e.g., the Fish Canyon Tuff and other monotonous
30 intermediates) give some insight into the potential magnitude of silicic LIP eruptions, as these
31 were erupted at the same time and along strike to the north of the ~0.4 Mkm³ Sierra Madre

1 Occidental silicic LIP, the largest ignimbrite-dominated volcanic province in North America
2 (Swanson et al., 2006; Ferrari et al., 2007). Several examples of similarly crystal-rich and
3 potentially large-volume, caldera-related rhyolitic ignimbrites occur within the Sierra Madre
4 Occidental (e.g., Copper Canyon and Vista Tuffs, Swanson et. al., 2006).

5 6 **5. Discussion**

7
8 Constraining the dimensions of LIP eruptive units are vital to determining potential rates of
9 magma eruption from vent systems, the duration of eruption and emplacement, and potentially,
10 the rates of magma production and storage (Tolan et al., 1989). This discussion focusses on these
11 aspects and examines what fundamental differences or similarities exist between the largest
12 basaltic and silicic eruptions.

13 14 *5.1 Eruption magnitudes of flood basalts vs rhyolites*

15
16 LIPs have been the site of the largest basaltic and silicic eruptions on Earth (Tables 2, 3). From
17 the available data on the dimensions of the products of individual eruptions, magma composition
18 does not appear to be a limitation to the upper size limits of erupted magma volume. However,
19 of the larger eruptions, most appear to be 1,000-3,000 km³, with a few enormous eruptions of
20 5,000 to possibly $\geq 10,000$ km³. It is worth noting that given the uncertainty of volume
21 estimations for the silicic eruptive units and the potential for considerable volumes to reside as
22 co-ignimbrite ash deposits, volume estimates for silicic eruptive units in LIPs are likely to be
23 considerably underestimated (especially for the Paraná-Etendeka quartz latites). Despite this first-
24 order similarity in the volume of individual eruptions, the basaltic and silicic eruptions are very
25 different in terms of vent types, eruptive mechanisms, magma discharge rates, eruption durations,
26 and emplacement styles (Table 4).

27
28 An important consequence for both basaltic and silicic eruptions is that the large magma volume
29 erupted has resulted in eruptive units with long run-out lengths and areal extents. The possible
30 existence of mafic long lava flows up to 1,000 km in the Deccan (Self et al., 2008) and on other
31 planets (Kezthelyi et al., 2006) require a large volume of mafic magma in order to satisfy mass

1 continuity, a condition only to be found in LIP eruptions (Self et al., 2008). Likewise, long run-
2 out lengths for pyroclastic density currents emplacing silicic ignimbrite are strongly dependent
3 on a high mass-flux to feed pyroclastic fountaining eruptions (Bursik & Woods 1996; Freundt,
4 1999; Branney & Kokelaar, 2002).

5 6 *5.2 Discharge rates*

7
8 The two very different vent configurations and eruptive styles (Table 4) play a significant role
9 in the orders of magnitude difference in discharge rates for the basaltic and silicic eruptions in
10 LIPs. For the flood basalt eruptions, early workers inferred fast eruption rates to emplace the
11 huge volumes of magma in a matter of days to a few weeks (Shaw & Swanson, 1970; Swanson
12 et al., 1975; Mangan et al., 1986; Tolan et al., 1989). Determining the discharge rate has obvious
13 implications in terms of the requirement for vast plumbing systems and magma reservoirs for the
14 magma to be delivered quickly to the surface. The identification of lengthy fissure systems of 70-
15 200 km long in the CRB Province (e.g., Swanson et al., 1975) provided the mechanism for the
16 rapid evacuation of a huge magma reservoir with the constraint that the mass discharge rate was
17 not so high as to generate a large eruption column. However, the vent system for only one
18 moderate-sized flood basalt eruption (Roza) has been studied in detail, and the nature and extent
19 of activity along other fissure vent systems and dykes for other flood basalt flow fields remain
20 poorly known.

21
22 The studies of Self and coworkers on the Roza flood basalt eruption have demonstrated that the
23 eruptive volume can be accounted for by an ~10 year duration at an averaged effusion rate of
24 ~4,000 m³ s⁻¹ (1.12 x 10⁷ kg s⁻¹), equivalent to the peak rate of the 1783-1784 Laki basaltic
25 eruption in Iceland (Self et al., 1997). However, given the ~150 km length of the total fissure
26 system for the Rosa flow, at these effusion rates, only part of the fissure system can have been
27 active at any one time. The eruption rate, if averaged over the fissure length, would be
28 exceedingly low (~0.0267 m³/s/m length of fissure) and result in magma freezing in the
29 dykes/fissures in transit to the surface. Such magma freezing may be one mechanism by which
30 effusion becomes concentrated or localised along the fissure system. Voluminous sheet lava
31 flows, therefore, do not require the rapid extrusion of mafic magma at rates much higher than in

1 historic eruptions except for the few cases of young flood volcanism, nor do long fissures imply
2 high eruption rates as only segments of the fissure may be active at any one time (cf. Hooper et
3 al., 2007).

4
5 Many flood basalt lavas in LIPs, were emplaced as inflated pahoehoe sheet flows at effusion rates
6 of 10^3 - 10^4 $\text{m}^3 \text{s}^{-1}$ ($\sim 2.7 \times 10^6$ - 2.7×10^7 kg s^{-1}). Some flood basalt lavas occur as rubbly pahoehoe
7 flows (e.g., Self et al., 1997; 1998; Kezsthelyi, 2002) where short-lived pulses at higher effusion
8 rates, potentially up to 10^6 $\text{m}^3 \text{s}^{-1}$, disrupted the thick, insulating crust and prevented the more
9 commonplace pahoehoe mode of emplacement (Kezsthelyi et al., 2006). One explanation is that
10 these surges of higher effusion rate may reflect enhanced periods of magma effusive activity over
11 greater lengths of the fissure system (Brown & Self, 2008).

12
13 The potential similarity in terms of discharge rate and therefore eruption style between small-
14 volume historic ($\leq 20 \text{ km}^3$; e.g., Laki) and flood basalt ($\geq 1,000 \text{ km}^3$) eruptions raise the possibility
15 that there is a physical limitation to how much, and how quickly, basaltic magmas can be
16 erupted. Despite the length of fissure vents, the effusion rates appear supply rate-constrained,
17 resulting in eruptions that must last years to 10's years to empty any stored magma reservoirs
18 despite their size. An insight into why flood basalt eruptions do not have extraordinarily high
19 discharge rates given the huge volumes of magma erupted may come from temporal
20 compositional trends observed in smaller volume intraplate basalt eruptions (Reiners, 2002). In
21 these cases, rates of melt extraction from the mantle and source-to-surface melt velocities of
22 $\sim 10^0$ - 10^1 km/yr are indicated, and may provide one limiting constraint on effusion rate at the vent
23 for the most primitive flood basaltic magmas with little evidence for crustal storage, assimilation
24 or fractionation. However, most flood basalt magmas require significant residence at crustal
25 levels, given their compositions are not primitive (e.g., Cox, 1980), and buoyancy forces or high
26 magma densities represent the main constraint on chamber evacuation and eruption rates.

27
28 Few constraints are available for discharge rates of silicic eruptions during LIP events. This is
29 in part due to the lack of information on the nature of source vents and the virtual absence of
30 plinian fall deposits in LIPs from which most constraints on magma discharge rates for silicic
31 explosive eruptions are made (e.g., Wilson et al., 1978; Sparks, 1986; Wilson & Walker, 1987;

1 Carey & Sigurdsson, 1989). The large-volume silicic eruptive units in LIPs are dominantly
2 ignimbrite or rheoignimbrite (e.g., Milner et al., 1992; Bryan et al. 2002; Ukstins Peate et al.
3 2005; Bryan, 2007) and the general lack of widespread plinian fall deposits suggest that in
4 general, mass discharge rates are sufficiently high ($>10^8$ - 10^9 kg/s) to prevent stable and buoyant
5 plinian eruption columns forming at the onset of, and during eruptions. Other factors that would
6 contribute to co-current eruption dynamics include large, wide or multiple vents, such as fissures
7 or along ring faults, and lower gas content (particularly for the rheomorphic ignimbrite examples)
8 that in turn help lower eruption velocity (Wilson et al., 1980; Woods, 1995; Freundt, 1999). In
9 general, the formation of single cooling units and in some cases, absence of internal erosion
10 surfaces or sedimentary deposits produced by epiclastic processes within ignimbrite sheets have
11 been interpreted to indicate ignimbrite emplacement within a period of no more than a few hours
12 or days (e.g., Christiansen, 1979; 2001).

13
14 The inferred intensities for a number of prehistoric and historic plinian eruptions vary between
15 1.6×10^6 kg s⁻¹ to 1.1×10^9 kg s⁻¹ (Carey & Sigurdsson, 1989), whereas the magma discharge rate
16 for the June 15, 1991 eruption of Mt Pinatubo was between 4×10^8 and 2×10^9 kg s⁻¹
17 (Koyaguchi, 1996). Assuming a discharge rate of 1×10^9 kg s⁻¹ and given the deposit volumes
18 of the large silicic eruptive units in Table 3, the estimated duration of these eruptions varies from
19 ~44 days for the smallest volume ignimbrite listed (Sana'a Ignimbrite, Afro-Arabian LIP) to ~248
20 days for the Guarapuava-Tamarana/Sarusas Quartz Latite (Paraná-Etendeka LIP). If mass
21 eruption rates were an order of magnitude lower, then eruption durations would be in the order
22 of 1-10 years, and approach the inferred durations of the flood basalt lava eruptions. The duration
23 of the first pulse of the Woongarra Rhyolite sill emplacement has been estimated at ~240 years,
24 equating to a magma discharge rate of 5×10^6 kg s⁻¹ (Trendall, 1995), which approaches
25 discharge rates of Quaternary sub-plinian eruptions (Carey & Sigurdsson, 1989). In such long-
26 lived eruptions, we would expect to see unsteadiness in the eruption that would be reflected in
27 bedding or multiple eruptive units in the deposits, which are rarely observed (cf. Sarusas Quartz
28 Latite, Ewart et al., 2004b). Alternatively, if eruption durations approach those from well-
29 documented Quaternary eruptions (ie. hours to days), then eruption intensities of 10^{10} to 10^{11} kg
30 s⁻¹ are required. Such high eruption intensities without the development of a tall Plinian eruption
31 column can be achieved by multiple vents or ring fracture fissure eruptions. The simple deposit

1 structures of the silicic eruptive units thus supports the notion for high eruptive fluxes (10^9 - 10^{11}
2 kg s^{-1}) and short duration (<1 month) eruptions. These higher rates are supported by recent work
3 on giant ash clouds (Baines & Sparks, 2005) that suggest eruption intensities approach 10^{10} kg
4 s^{-1} resulting in durations between 2-10 days for the largest eruptions (M8-9). Additional
5 supporting examples for short durations include the 450 km^3 Bishop Tuff eruption from Long
6 Valley, California, about 770,000 years ago that has been estimated to have lasted about 4 days
7 (Wilson & Hildreth 1997), and the compositionally zoned $2,200 \text{ km}^3$ Huckleberry Ridge Tuff
8 erupted from the Yellowstone volcanic field, which occurs as a single cooling unit and likewise
9 been suggested to result from an eruption that took days at most (Christiansen, 1979, 2001).

11 *5.3 Frequency of large magnitude (>M8) eruptions from LIPs*

12
13 It is both the volume of magma emitted during individual eruptions in LIP events and the total
14 volume of magma released (>0.1-80 Mkm^3) that make LIP events so exceptional in Earth history
15 (Self et al., 2005; Self, 2006). It is this combination of large erupted volumes and high frequency
16 that lead to the rapid construction of thick (1->3 km) areally extensive plateaus (0.1 - 2 Mkm^2),
17 which internally, show few signs of major time breaks, erosion surfaces and regional
18 unconformities (Fig. 6A). The high-frequency of large-magnitude eruptions also distinguishes
19 LIP events from other tectonic settings and processes where igneous rocks are formed.
20 Importantly, without LIP-forming igneous events, basalt super-eruptions would not have occurred
21 through Earth history, but in contrast, silicic super-eruptions have occurred independently of LIP
22 events (Sparks et al., 2005).

23
24 From a volcanological viewpoint, geochronological studies of LIPs, summarized in studies such
25 as Rampino & Stothers (1988), Courtillot & Renne (2003), Kelley (2007), Bryan & Ernst (2008)
26 and Chenet et al. (2008), have revealed two main features relevant to the timing of eruptions and
27 LIP formation. These are that: 1) much (70-90%) of the eruptions are produced during one or two
28 main pulses of eruptive activity; and 2) that the pulse or pulses, or even the whole duration of
29 activity in the LIP, can be very brief geologically, <5 Ma, and possibly even < 1 Ma in some
30 cases. For LIPs of any age, the errors on the age estimates cannot resolve individual formations
31 within these pulses of activity, and certainly cannot resolve individual eruptions. In many cases,

1 the errors encompass the age range of almost all eruptive units from top to bottom of the LIP pile
2 (e.g., Barry et al., 2010).

3
4 For the flood basalt eruptions, evidence from the CRB LIP suggests that during the main pulse
5 and emplacement of the Grande Ronde and Wanapum Basalt Formations (~16-5-15.3 Ma), a
6 high frequency of the largest magnitude eruptions existed producing the most voluminous flow
7 fields (Tolan et al., 1989). For example, the Grande Ronde Basalt forming >60% of the total
8 volume of the CRB LIP (Camp et al., 2003), comprises at least 110 individual eruption packages
9 or flow fields (Barry et al., 2010) with an averaged volume of 1,238 km³ (Tolan et al., 1989).
10 New ⁴⁰Ar/³⁹Ar dates for Grande Ronde lavas reveal they were emplaced within a maximum time
11 range of 0.42 ± 0.18 Myr (Barry et al., 2010), corresponding to an averaged frequency of ≥M8
12 eruptions of 220/Myr or one ≥M8 eruption every ~4,200 yrs. For larger continental flood basalt
13 provinces such as the Deccan, recent studies (Self, 2006; Self et al., 2008; Chenet et al., 2009)
14 indicate that individual formations emplaced over similar time scales of hundreds of thousands
15 of years, have volumes either equivalent to the main phase lavas (Grande Ronde Formation,
16 ~0.15 Mkm³) or to the entire CRB LIP (~0.23 Mkm³, Camp et al., 2003). Therefore, LIP
17 formations must also be characterised by even higher frequencies of M8 and larger eruptions.

18
19 For most LIPs other than the CRB, however, the number of eruptions is unknown. Even within
20 the CRB, the number is not known precisely but is probably around 200. Fitting the number of
21 eruptions within a 1-2 Ma timeframe still gives average eruption intervals of 1,000s - 10,000 yrs
22 (Self et al., 2006; Barry et al., 2010). While the accumulated lava pile, and the thickness added
23 by each eruption, is impressive, considered from a modern or historic perspective, LIP eruptions
24 probably were not necessarily 'hyperactive', even during the main pulse. Much more needs to
25 be known about the rates of lava production and the lengths of hiatuses between eruptions in
26 LIPs, and various approaches are being undertaken to estimate this (Chenet et al., 2008; Jolley
27 et al., 2008), but such studies are in their infancy.

28
29 The recent compilation on large-volume silicic explosive eruptions by Mason et al. (2004)
30 revealed 42 known eruptions of >M8 over the past 36 Ma. This yielded a minimum time-
31 averaged estimate of eruption frequency of 1.1 events/Myr since the beginning of the Oligocene,

1 but over this time, such large magnitude eruptions have clustered in two pulses at 36-25 Ma and
2 13.5 Ma to present. The older pulse corresponds to two LIP events: the Afro-Arabian and Sierra
3 Madre Occidental provinces but with most dimensional data for eruptions drawn from related
4 large-volume silicic ignimbrite volcanism to the north of the Sierra Madre Occidental in the
5 Great Basin region of western U.S.A. (e.g., Gans et al., 1989; Best & Christiansen, 1991). During
6 these pulses, large eruption frequencies were slightly higher at ~2 events/Myr. However, virtually
7 no data are currently available on the magnitude of individual silicic eruptions within the Sierra
8 Madre Occidental, and eruption frequencies will have been much higher for this period. In
9 contrast to the eruptive record for the last 35 Myrs, large eruption frequencies were at least 9
10 events/Myr during the Paraná-Etendeka LIP event, and for the Afro-Arabian province, the
11 equivalent of 12 events/1 Myr. Consequently, the frequency of silicic super-eruptions is greater
12 during LIP events than when compared to global, long-term averaged frequencies of silicic super-
13 eruptions. Importantly though, based on available age data for the continental flood basalt
14 provinces, these high frequencies are sustained only for very brief periods of ≤ 1 Myr.

15 16 *5.4 Generation & storage of large magnitude LIP eruptions*

17
18 The eruption of such exceptionally voluminous magmas and the often remarkable chemical
19 homogeneity in on-land deposits have led to the general interpretation that LIP eruptions require
20 very large magma reservoirs, which are rapidly evacuated. The very large volumes for individual
21 eruptions from LIPs raise a number of space-volume issues in terms of the storage and
22 dimensions of, and connectivity and interactions between holding chambers within the crust. As
23 noted by Ewart et al. (1998b), the erupted volume of the Springbok Quartz Latite equates to a
24 magma sphere diameter of 23 km. Alternatively, a sill-like magma body with dimensions of ~4
25 x 40 x 40 km also approximates the minimum erupted volume of the Springbok Quartz Latite.
26 The Messum igneous complex, identified as the eruptive centre for the Springbok Quartz Latite
27 (Milner & Ewart, 1989) is a roughly circular structure of ~18 km diameter, but characterised by
28 downsagged margins and inward-dipping strata following the quartz latite eruptions (Ewart et
29 al., 2002). Volumetric considerations suggest that chamber dimensions were much greater than
30 the collapse diameter of Messum, hence resulting in regional downsagging rather than classical
31 Valles-type caldera collapse structures (see also Volcan de l'Androy, Madagascar LIP; Mahoney

1 et al., 2008). The lack of well-defined calderas in LIPs may thus be in part due to the greater
2 depth and lateral dimensions of the magma chambers.

3
4 Not only do issues of space issues arise from the storage of such huge volume magma bodies in
5 the crust, but additional complications arise in a temporal sense because many >M8 basaltic and
6 silicic eruptions may occur within 1 Myrs, some or all of which may be genetically unrelated and
7 represent new episodes of large-volume magma generation. For example, a number of large
8 volume (>1,000 km³), closely-spaced ($\leq 10^5$ yrs) and genetically related silicic eruptions
9 (Goboboseb and Springbok Quartz Latites) occurred from the Messum igneous complex, and
10 which overlapped in space and time with flood basaltic eruptions (Ewart et al., 1998b; 2002).
11 How do these magma reservoirs spatially overlap and how does this impact on the thermal and
12 rheological character (e.g., de Silva & Gosnold, 2007) of the crust? The architecture and spatial-
13 temporal relationships of flood basaltic and rhyolitic magma reservoirs in the crust during LIP
14 events remain poorly understood.

15
16 Considerable petrographic and chemical evidence indicates that many flood basaltic magmas
17 underwent storage, fractional crystallisation and crustal assimilation at lower crustal pressures
18 prior to eruption (Cox, 1980; Hooper, 1997; Ewart et al., 1998a; Ramos et al., 2005; Hooper et
19 al., 2007), as well as potentially further storage, fractionation and degassing at shallow depths
20 in sill complexes prior to final extrusion (e.g., Cox, 1980; Rodriguez Durrand & Sen, 2004;
21 Puffer et al., 2009). However, the lack of caldera collapse structures associated with flood basalt
22 magma withdrawal implies that any reservoirs were most likely located at middle to lower crustal
23 depths (Hooper, 1997) or that magma is both intruded into and inflating, rather than evacuating
24 shallow-level sills while also being extruded at the surface (Puffer et al., 2009). In contrast, some
25 flood basalt lavas show little evidence for crustal storage or assimilation and retain primitive
26 mantle geochemical and isotopic signatures (e.g., Tafelkop and Santa Lucia basalts, Paraná-
27 Etendeka Province, Ewart et al., 1998a; Kirstein et al., 2000). Although these magmas can be
28 interlayered with the more common fractionated basaltic magmas, they must have a different
29 transport history and pathway from mantle source regions to the surface.

30
31 While LIPs often show evidence for bimodality and a prominent silica or Daly gap, recent studies

1 of the Paraná-Etendeka and Afro-Arabian LIPs have revealed the presence of more intermediate
2 magma compositions (Ewart et al., 2004b; Ukstins Peate et al., 2008) that reduce the significance
3 of the compositional gap between the associated flood basalts and rhyolites. For the Paraná-
4 Etendeka province, intermediate (latite) compositions are small in volume and restricted in
5 distribution that have hindered their recognition, but they are significant in establishing
6 petrogenetic linkages between the associated flood basalt and rhyolitic magmas (e.g., Ewart et
7 al., 2004b). In contrast, it is in distal tephra records to the Afro-Arabian LIP that a more complete
8 compositional spectrum of LIP magmas is recorded. Tephra from the Indian Ocean (~3,000 km
9 away), which have been temporally, geochemically, isotopically, and paleomagnetically
10 correlated to individual silicic eruptions in Yemen and Ethiopia, preserve a complete spectrum
11 of compositions ranging from basalt (43 wt% SiO₂) to rhyolite (75 wt% SiO₂) in individual
12 tephra shards from within a single eruption (Ukstins Peate et al., 2008). In addition, banded
13 shards from individual eruptions record up to 85 % of the SiO₂ variation observed in the entire
14 eruption at the scale of <1 mm³, clearly demonstrating that the entire magmatic compositional
15 variation found within a LIP can potentially be sampled by individual eruptive events. This
16 requires that the full spectrum of compositions, representing an extreme range from basaltic
17 through silicic, is present within magma chambers at the time of eruption, and argues for magma
18 chamber connectivity, efficient evacuation and large as well as small-scale mingling and mixing
19 during eruption.

20
21 Many documented large-volume (>1,000 km³) silicic eruptions in the recent geologic record are
22 examples of crystal-rich (30-50%) dacite-rhyolites labelled as “monotonous intermediates” by
23 Hildreth (1981), including the four of the largest eruptions occurring since 35 Ma listed by
24 Mason et al. (2004): the Fish Canyon Tuff (an archetypal monotinous intermediate), Toba Tuff,
25 Lund Tuff and Atana Ignimbrite. Recent studies of the Fish Canyon Tuff by Bachmann et al.
26 (2002, 2005) recognised that this and other crystal-rich silicic ignimbrites record the eruption of
27 batholithic volumes of crystal mush forming in upper crustal magma chambers. Whereas the
28 rejuvenated batholith model (e.g., Bachmann & Bergantz, 2003, 2006, 2009) satisfactorily
29 explains the essential features of the large-volume “monotonous intermediates”, this genetic
30 model cannot be applied to the similarly large and larger-volume rhyolite eruptions from the
31 continental flood basalt provinces. Voluminous deposits of crystal-rich ignimbrites (i.e.

1 monotonous intermediates) for example, are absent from the continental flood basalt provinces,
2 and Streck & Grunder (2008) have provided other arguments against mush extraction models for
3 high-temperature, crystal-poor Fe-rich rhyolites. Many LIP rhyolites are distinguished by their
4 large volume, crystal-poor nature, anhydrous mineralogy, high crystallisation and eruption
5 temperatures ($>1,000^{\circ}\text{C}$) and strong geochemical evidence for a petrogenetic link with associated
6 flood basalts (Bellieni et al., 1986; Garland et al., 1995; Peate, 1997; Ewart et al., 1998b; 2004b;
7 Ukstins Peate et al., 2008). An alternative model for the origin of these large-volume crystal-poor
8 LIP rhyolites is thus required.

9
10 We consider that the variety of large-volume basaltic and silicic eruptions in LIPs can be
11 envisaged in terms of four end-member magma petrogenetic pathways (Fig. 10). Certain flood
12 basaltic magmas preserve mantle geochemical signatures and show little to no evidence for
13 storage at crustal depths, low pressure crystallisation and crustal assimilation. These smaller
14 volume flood basalt magmas may have been transferred directly from melting regions in the
15 upper mantle to fissure vents at surface or have resided temporarily in reservoirs in the upper
16 mantle or in mafic underplate, thereby preventing opportunities for crustal contamination
17 (pathway A in Figure 10). In contrast, it is more common for the larger volume flood basalt
18 magmas (notably low-Ti types) to have undergone storage at lower \pm upper crustal depths
19 resulting in crustal assimilation, crystallisation, and degassing to produce aphyric to plagioclase-
20 dominant basaltic and basaltic andesite lavas (pathway B1 in Figure 10). Large volume high-Ti-
21 type flood basalts (e.g., Khumib-Urubici-type, Ewart et al., 2004a) also show evidence for crustal
22 storage and shallow-level fractionation prior to eruption but have not undergone (silicic) crustal
23 assimilation (pathway B2 in Figure 10).

24
25 For the crystal-poor large-volume rhyolites in continental LIPs, most workers have invoked large-
26 scale AFC processes involving lower crustal granulite melting and/or remelting of basaltic
27 underplate with additional mafic and silicic magma inputs and extended fractional crystallisation
28 (pathway C in Figure 10) to generate the high-temperature anhydrous, crystal-poor silicic
29 magmas characteristic of LIPs such as the Paraná-Etendeka and Karoo (e.g., Piccirillo et al.,
30 1987; Garland et al., 1995; Ewart et al., 1998b; 2004b; Miller & Harris, 2007). Surface
31 expressions of these eruptions may be regional sag structures (Ewart et al., 2002) or fissure vents,

1 with well-defined surface calderas absent due to high-temperature ($>1,000^{\circ}\text{C}$) magma reservoirs
2 residing at deeper crustal levels. Chambers at mid to lower crustal levels are more likely to lack
3 chilled envelopes or sidewall crystallisation and undergo strong convective mixing promoting
4 chemical homogenisation.

5
6 Large-volume, lower temperature ($<900^{\circ}\text{C}$) crystal-rich rhyolite ignimbrites, as found in the
7 silicic LIPs (Bryan, 2007) and extensional continental margin environments (e.g., Altiplano Puna
8 province), appear to have an origin similar to that described for the Fish Canyon Tuff and other
9 ‘monotonous intermediate’ ignimbrites. In these cases, the eruptions reflect remobilisation of
10 near-solidus magmatic mushes or magma bodies of several thousands of km^3 (with $\sim 45\text{-}50$ vol.%
11 crystals) that are in the process of forming upper-crustal batholiths (pathway D1 in Figure 10).
12 Shallow intrusion and underplating by flood basaltic mafic magmas, which can be enhanced by
13 active extension (Ferrari et al., 2009) provide the necessary thermal and volatile inputs to trigger
14 the eruption of these upper crustal chambers and the development of well-defined calderas up
15 to 80 km in diameter. By contrast, upper crustal intrusion of basalt also leads to the rapid
16 generation and eruption of moderate volumes of crystal-poor rhyolite through the remelting of
17 solidified and highly differentiated plutonic rocks formed during preceding phases of magmatism
18 in the LIP event (pathway D2 of Figure 10), and the Alacrán ignimbrite (Bryan et al., 2008) in
19 the Sierra Madre Occidental silicic LIP is an example of this petrogenetic pathway.

20 21 **6. Conclusions**

22
23 Large igneous provinces have been the loci for both basaltic and silicic super-eruptions ($>M8$)
24 throughout Earth history, and are therefore important for understanding their potential for driving
25 environmental and climate change and causing mass extinctions, melt production rates from the
26 sublithospheric mantle and crust, the thermal, mechanical and compositional evolution of the
27 lithosphere, and what upper limits there may be to the volume and rate of magma eruption. LIPs
28 are unique for their substantial cumulative volumes ($>10^5\text{-}10^7 \text{ km}^3$) of emplaced magma over
29 brief periods (1-5 Myrs) that ultimately results from 10's to 100's of $>M8$ eruptions and
30 intrusions. The compilation here indicates that basaltic and silicic super-eruptions from LIPs have
31 similar magnitudes, but the difficulties in identifying co-ignimbrite ashes for the silicic eruptive

1 units result in present volumes being significant underestimates. Nevertheless, magma
2 composition appears to be no barrier to the volume of magma erupted, and several basaltic and
3 silicic super-eruptions from LIPs are estimated to have been >M9.2, and consistent with a
4 calculated upper limit to eruption size by Mason et al. (2004). Based on the data set presented
5 here and that of Mason et al. (2004), it appears that an upper limit of ~M9 exists for other
6 tectonic settings, but for LIPs a more realistic upper limit of eruption magnitude is closer to M9.5
7 (or ~11,000 to 12,000 km³ of magma). Supply rate and duration seem to be the primary factors
8 controlling the run-out length, areal extent and volume of the basaltic and silicic eruptive units.
9 These basaltic and silicic super-eruptions do significantly differ in terms of eruptive frequency,
10 sources, style and emplacement mechanisms, but have petrogenetic linkages. The dominantly
11 effusive nature and lower discharge rate (10^6 - 10^8 kg s⁻¹) for the basaltic super-eruptions are
12 inferred to result in long durations (yrs to >10 yrs) for individual eruptions, which in some cases
13 may be a melt generation rate-limited process. In contrast, the deposit characteristics of silicic
14 super-eruptions in LIPs suggest very high-rate (10^9 - 10^{11} kg s⁻¹), short-lived eruptions, similar to
15 eruption durations of hours to days for well-studied Quaternary eruptions; such eruptive rates
16 require the availability of 1,000's km³ of silicic magma stored in crustal chambers to be erupted.
17 Importantly, detailed studies of LIPs are revealing evidence for the contemporaneous interaction
18 and eruption of basaltic and silicic magmas, which challenges concepts of how and where in the
19 lithosphere such large volumes of magma are stored and what pathways they utilise to erupt at
20 the surface. Further studies are required to rigorously establish eruptive unit correlations for all
21 LIPs, and linking onshore silicic eruptive units with distal and marine -deposited co-ignimbrite
22 ash deposits that will give considerable insight into the eruptive mechanisms, global ash
23 dispersion (e.g., Peate, 2009), and environmental effects from individual or successive, closely-
24 spaced large magnitude (>M8) super-eruptions.

25

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1 **Figures**

2

3 Figure 1. Examples of intervening deposits recording time breaks between basaltic LIP eruptions.

4 A) Gently folded Columbia River Basalt lavas exposed at Sentinel Gap (46° 48.750' N 119°

5 56.910' W), Washington, USA, showing part of the younger Wanapum Formation lavas over

6 lying upper part of the voluminous Grande Ronde Formation lavas and separated by the Vantage

7 sediment inter-bed (white layer, up to 20 m thick). Total height of cliff is ~ 300 m. B) Contact

8 between lavas of the Ambenali Formation of the Wai Sub-group (Deccan LIP) with a massive

9 base of the upper lava's core (note, virtually no lower crustal zone) resting on a mechanically

10 eroded/brecciated S-type pahoehoe top of upper crustal zone of lower flow with relatively thick

11 red-weathering zone of silty-clayey material between the breccia clasts. C) Aeolian cross-bedded

12 sandstone interbedded with olivine-phyric compound pahoehoe flood basaltic lavas of the

13 Paraná-Etendeka LIP, Huab Outliers, Namibia (20° 39.161' S 14° 09.350' E).

14

15 Figure 2. Examples of intervening deposits recording time breaks between silicic eruptions in

16 LIPs. A) Intracaldera lacustrine sedimentary succession developed between two major ignimbrite

17 and caldera-forming eruptions. This well-bedded, fine-grained low energy sedimentary package

18 overlies a capping lag breccia unit to a welded rhyolitic ignimbrite, and is itself overlain by a

19 flow banded rhyolite lava flow/dome complex with a marginal hyaloclastite and fluidised

20 sediment contact zone; Whitsunday Island (20° 18.545'S 149° 2.873'E), Whitsunday silicic LIP.

21 B) Normally graded, clast-supported and matrix-poor, resedimented pyroclastic unit overlying

22 a reddened top to a fine-grained rhyolitic ignimbrite at left and capped by a welded, fine-grained

23 base to a rhyolitic ignimbrite at right. Contacts marked by yellow dashed lines (21° 41.305' N

24 103° 52.520' W; San Martin de Bolaños mine, drill hole Z-425 @ 147.3- 141.4 m depth; Sierra

25 Madre Occidental silicic LIP, Mexico).

26

27 Figure 3. Textural features illustrating the synchronicity of mafic and silicic magmas in LIPs. A)

28 Mafic and silicic magma interaction producing 'organ-pipes' (21° 25.72' S 14°16.89' E), the

29 result of multiple diapirs of granite that have intruded upwards through unconsolidated mafic diorite

30 in the easternmost moat of the Messum igneous complex (Paraná-Etendeka LIP; see also Ewart

31 et al., 2002). B) Outsized, ductile deformed juvenile mafic spatter clast within the lava-like Jozini

1 rhyolite (Karoo LIP), Namaacha Falls, Mozambique ($25^{\circ} 57.55' S$ $32^{\circ} 02.1' E$). Pen is 15 cm
 2 long. C) Outsized, juvenile mafic scoria clast within a densely welded rhyolitic ignimbrite, with
 3 the eutaxitic fabric subhorizontal in photo (Whitsunday silicic LIP, Cid Island, northeast
 4 Queensland; $20^{\circ} 16.341'S$ $148^{\circ} 55.083'E$). Pen lid width is 0.8 cm.

5
 6 Figure 4. Potential examples of contemporaneous eruptions in LIPs. A) Local correlation of
 7 Urubici lava flows in the São Joaquim area (Peate et al., 1999) with local topographic relief
 8 controlling the emplacement of early lavas flow units (3-8). PE, Perico; AB, Aguas Brancas; UR,
 9 Urubici; SM, Morro do Igreja; CO, Corvo Branco; RR, Rio Rufino sections of Peate et al.(1999).
 10 B) Compositional plot of MgO-TiO₂ showing the similarity of basaltic lava units 9 & 10 and their
 11 difference to other Urubici high-Ti lava flow units. The basaltic lava interbedded with the
 12 Urubici units 9 & 10 in the UR profile is sample DUP-38 (from Peate & Hawkesworth, 1996)
 13 that is an evolved Gramado-type basalt with 3.3 wt% MgO and 1.68 wt% TiO₂. C) Simplified
 14 map of southern Etendeka showing the distribution of the Etendeka flood volcanic succession
 15 and schematic stratigraphic sections for the Awahab and Goboboseb Mountain area (modified
 16 from Jerram et al., 1999) where two compositionally different mafic lava units separate units 1
 17 and 2 of the Goboboseb Quartz Latite. In the Goboboseb Mountains, the Copper Valley
 18 icelandite (CVI, Ewart et al., 1998b) lies stratigraphically between Goboboseb units 1 and 2,
 19 whereas chemically different, low-Ti-type basalts separate these quartz latite units in the Awahab
 20 area. QL, quartz latite.

21
 22 Figure 5. Examples of large-volume silicic eruptive units from LIPs. A) View northeast towards
 23 Tafelkop ($21^{\circ} 07.13'S$ $14^{\circ} 17.3'E$) that is capped by two ~30 m sheet-like quartz latite
 24 rheoignimbrite units (Goboboseb I, Gb 1; Goboboseb II, Gb 2). B) Closeup of base of Goboboseb
 25 I in contact with brecciated basaltic lava at base illustrating a typical 'valley-fill' geometry. A
 26 basal, discontinuous topography-filling, massive quartz latite facies is overlain by an intensely
 27 platy jointed facies (PJF) that is more sheet-like in geometry (boundary between depositional
 28 facies marked by dashed line). Location of section shown in A). C) Afro-Arabian main silicic
 29 sequence of pyroclastic flow and fall deposits from Bayt Baws, located near Sana'a, Yemen (15°
 30 $16.5' N$ $44^{\circ} 11.3' E$). Main eruptive units, from base to top are: 1) Jabal Kura'a Ignimbrite; 2)
 31 Escarpment Ignimbrite; 3) Green Tuff; 4) SAM Ignimbrite; 5) caldera-collapse breccia Iftar

1 Alkalb; conglomerate separates the Escarpment Ignimbrite and Green Tuff (Ukstins Peate et al.,
2 2005). The ignimbrites exhibit distinct tabular morphology, which is recognizable across 100's
3 of km around the Sana'a basin and towards the western escarpment. Total thickness of section
4 is 115 m.

5
6 Figure 6. Examples of large-volume basaltic units from LIPs. A) Deccan lavas from the Wai
7 Sub-Group at Arthur Seat in the Western Ghats near Mahabaleshwar, India ($17^{\circ} 58.780' N 73^{\circ}$
8 $38.225' E$), showing several major basalt lava sheet lobes. The main parting in the middle of the
9 photo is the base of a 60 to 70-m-thick sheet lobe. Below this, three thinner (~20 m thick) sheet
10 lobes show sloping, vegetated upper crustal zones and cliff-forming lava cores. The core-upper
11 crust level in the thick sheet lobe is a horizon with overhangs, and the top of the lobe is the lower
12 of two small, prominent cliff lines. Above this is a vegetated slope capped by another cliff-line
13 (near top of photo), which is another flow-field composed of smaller (~100 m long x 10 m thick)
14 inflated pahoehoe lobes without thick cores. B) Overview of the Sand Hollow flood basalt flow
15 (Table 2) illustrating internal morphology of a single, large-magnitude, ~60 m thick sheet lobe.
16 An upper and lower zone of wider-spaced joints in core of lobe are separated by a more closely
17 spaced central jointed zone. Darker, slightly banded, and gullied part approximates the upper
18 crustal zone of the lobe (top removed by erosion); base is in vegetated slope. Palouse Falls (46°
19 $39.730' N 118^{\circ} 13.554' W$), Columbia River LIP, USA. C) Coastal exposure of basaltic lava flow
20 unit showing well-developed columnar jointing, near Trongisvágur, Suðuroy Island, Faroes
21 (North Atlantic LIP). D) Overview of the ~300 m thick Finger Mountain Sill of Jurassic Ferrar
22 Dolerite, Upper Taylor Glacier, Antarctica ($77^{\circ} 44.45' S 160^{\circ} 42.78' E$), intruded into the Beacon
23 Sandstone formation. Part of the Dry Valleys nested sill complex and plumbing system for the
24 Kirkpatrick flood basalts (Ferrar LIP).

25
26 Figure 7. Map of the pre-rift juxtaposition of South America and Africa showing the Paraná-
27 Etendeka large igneous province (modified from Peate et al., 1992 and Nardy et al., 2008) and
28 the areal extents of correlated silicic eruptive units discussed in text and listed in Table 3.

29
30 Figure 8. Discrimination diagrams, in part from Nardy et al. (2008) comparing the Sarusas and
31 Ventura quartz latites from the Etendeka province with the newly defined Guarapuava and

1 Tamarana silicic compositions from the Paraná province.

2
3 Figure 9. Enrichment-depletion diagrams showing the variation in major and trace elements for
4 correlated silicic units from the Paraná and Etendeka provinces. In A) the averaged composition
5 of the Etendeka Goboboseb II quartz latite is normalised to the averaged composition of Paraná
6 PAV-A rhyolite (Whittingham, 1991); and in B) the averaged composition of the Springbok
7 quartz latite is normalised to the averaged composition of PAV-B of Whittingham (1991). The
8 Goboboseb and Springbok quartz latites have most likely been erupted from the Messum igneous
9 complex (Ewart et al., 2002) in Namibia suggesting run-out distances of up to ~500 km into the
10 Paraná province. The diagram illustrates subtle geochemical variations between the correlated
11 units, interpreted to represent a proximal to distal chemical variation. For these different
12 eruptions, lateral geochemical variations are similar, with notable depletions in phenocryst-
13 compatible elements, such as Ti, Mn, P and Ca, suggesting that density segregation of
14 phenocrysts during flow over 100's km may have contributed to this chemical variation and
15 produced relatively crystal-poor and ash-enriched deposits in the Paraná. Geochemical data for
16 the Goboboseb II and Springbok quartz latites are from Milner (1988) and Ewart et al. (1998b);
17 data for the PAV-A and PAV-B are from Whittingham (1991); n is number of analyses on which
18 averaged compositions are based.

19
20 Figure 10. Conceptual crustal view (not to horizontal scale) of four end-member petrogenetic
21 pathways for large magnitude (basaltic and silicic) eruptions, principally in continental LIPs. The
22 effects of crustal thinning on magma generation are not included in this depiction. Sills of
23 mantle-derived basaltic magma are injected principally near the crust-mantle boundary. Over
24 time this underplated basaltic magma produces seismic (top of basaltic underplate) and petrologic
25 (base of underplate) mohos. Basaltic magmas may be either (A) extracted rapidly from melting
26 source regions in the mantle and erupted at the surface or (B) pond in lower crustal magma
27 chambers where the magmas become subject to open system processes (assimilation, magma
28 mixing, melt extraction) as well as fractional crystallisation. Low-Ti-type flood basalt magmas
29 (B1; e.g., Tafelberg-Gramado lavas, Ewart et al., 1998a) have typically experienced assimilation
30 of intermediate to silicic composition lower crust, whereas high-Ti-type flood basalt magmas
31 (B2; e.g., Khumib-Urubici lavas, Ewart et al., 2004a) have typically experienced magma mixing

1 ± remelting of the newly formed underplate. In both cases, additional upper crustal storage (B3)
2 can result in further crustal assimilation, crystallisation and magma degassing. High-temperature
3 (>950°C) silicic magmas are genetically related to the low- and high-Ti type flood basaltic
4 magmas and large-scale crustal assimilation characterises low-Ti (C1) silicic magmas (e.g.,
5 Goboboseb and Springbok Quartz Latites, Ewart et al., 1998b), whereas high-Ti silicic magmas
6 (C2) show little to no evidence for the involvement of silicic crust in their petrogenesis (e.g.,
7 Chapecó-type rhyolites, Garland et al., 1995; Ewart et al., 2004b), and appear to be rapidly
8 extracted from magma chambers residing in the lower crust. Thermal and mass fluxes of basalts
9 to mid to upper crustal depths result in either the remobilisation of felsic cumulate piles or
10 partially crystallised batholiths (D1) akin to the model for the Fish Canyon Tuff proposed by
11 Bachmann & Bergantz (2002). Additionally, remelting of differentiated and solidified granitic
12 intrusions (D2, e.g., Alacrán Tuff of the Sierra Madre Occidental, Bryan et al., 2008) can also
13 occur, producing moderate to large volumes of relatively low-temperature (<850°C) high-K₂O
14 rhyolites.

15

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